



# Uncertainty and Underestimation of GRB Prompt Emission Duration with Swift/BAT

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#### **T90 Distribution**



Kouveliotou et al., 1993

#### **T90 Distribution - Instrumental Effects**



Lien A. et al., 2016

#### **Project Goal**

To determine if the measured T90 durations of GRBs observed with *Swift*/BAT are accurate approximations of their true durations

#### Instrumental Considerations- SVOM/Eclairs







Fig. 16 Detection fraction (the ratio between the number of positions at which the burst is detected and the number of simulated positions) for each of the 10 ulGRBs according to the redshift.

Dagoneau, Schanne, Atteia, Götz, and Cordier (2020)

#### True Light Curve









## The Neil Gehrels Swift Observatory Burst Alert Telescope





<u> https://www.nasa.gov/mission\_pages/swift/spacecraft/</u> https://swift.gsfc.nasa.gov/about\_swift/bat\_desc.html

# *Swift*/BAT Instrumental Parameters

#### Relevant Parameters:

- Number of Active Detectors (NDETS)
- Incident angle (PCODE)
- Background
- Not Relevant:
  - Energy band



#### Lien A. et al., 2014



# *Swift*/BAT Instrumental Parameters

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Lien A. et al., 2014

### PCODE and Incident Angle



Incident Angle = 0 deg PCODE = 1



#### **PCODE and Incident Angle**



Incident Angle = 29 deg PCODE = 0.5



#### **PCODE** and Incident Angle





#### PCODE vs BAT Field of View

Not exactly one-to-one with incident angle



# *Swift*/BAT Instrumental Parameters

#### **Relevant Parameters:**

- Number of Active Detectors (NDETS)
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  - Energy band



#### Simulation Method Summary

Template Light Curve



Simulate Source Light Curve

Select T100 signal Fold with BAT response matrix Add flat background

### Simulation Method Summary



### **Simulation Method Summary**



















#### GRB090510

















#### **Distance Considerations**



Littlejohns et al., 2013

#### Kocevski and Petrosian, 2013

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#### **Distance Considerations - Results**

This is preliminary, to be further investigated at a later time.



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#### Conclusions

T90 may not be an accurate representation of true source duration due to instrumental effects

The T90 is not enough to distinguish between Short and Long GRBs May be possible to quantify (and correct for) these biases

How are the observed luminosity and T90 distributions affected? How do other instruments behave?

# Thank you!

# Backup Slides

### GRB Light Curves FRED-like





# GRB Light Curves Symmetric





# GRB Light Curves Unique Selection







### Instrument Considerations









Lien A. et al., 2016

Frontera et al., 2009

Qin et al., 2013