

# Instrumental Tip-of-The-Iceberg Effects on the Prompt Emission of *Swift*/BAT GRBs

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[mikejmoss3@gmail.com]

Arxiv: <https://arxiv.org/abs/2111.13392>

Collaborators:

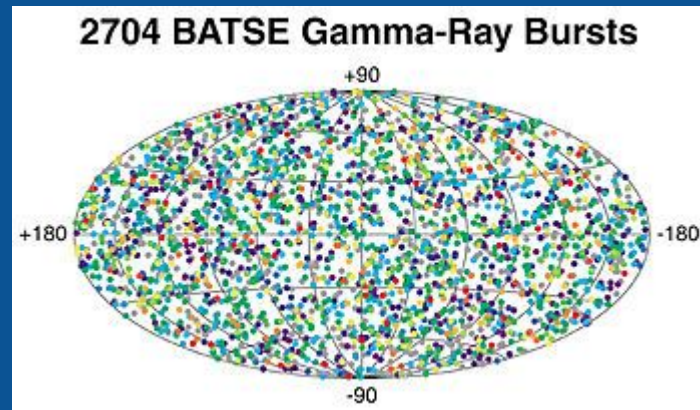
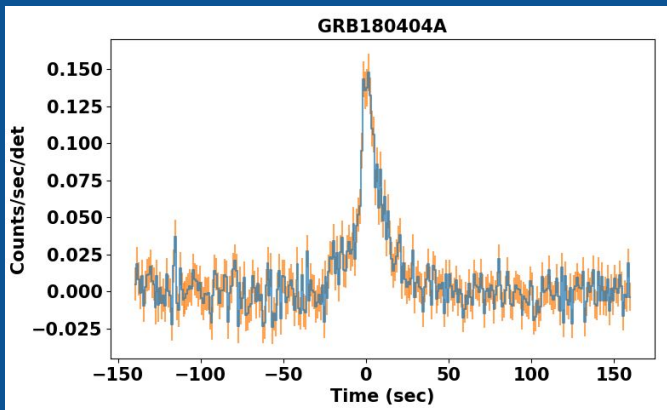
Amy Lien (NASA/GSFC, CRESSTII, UMBC),

Sylvain Guiriec (NASA/GSFC, CRESSTII, GWU)

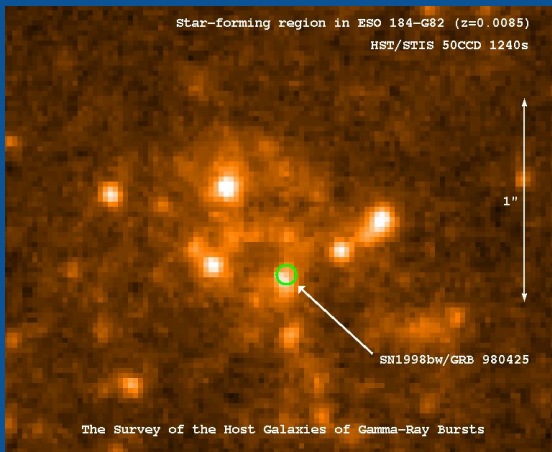
Brad S. Cenko (NASA/GSFC, UMD)

Takanori Sakamoto (AGU)

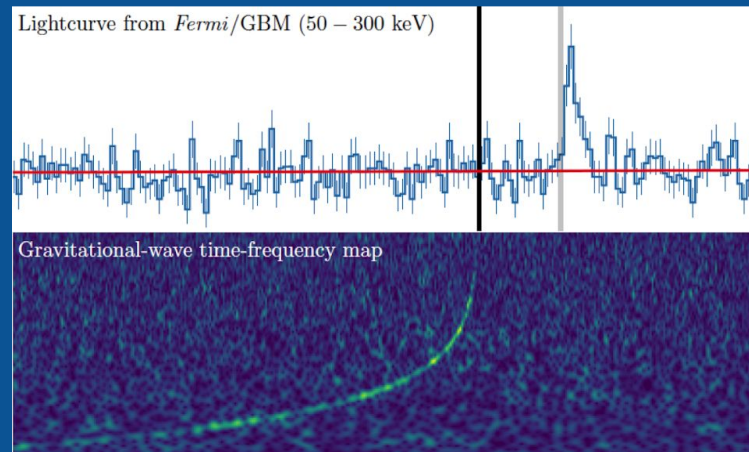
# Context - GRB Phenomena



Meegan et al., 1992

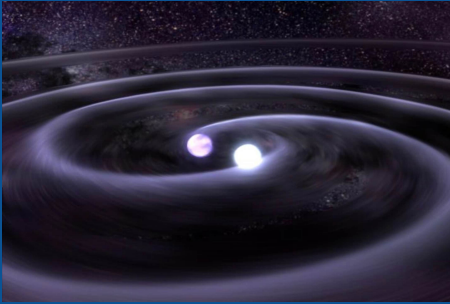


Holland et al., GCN 704



[LivingLigo post](#)

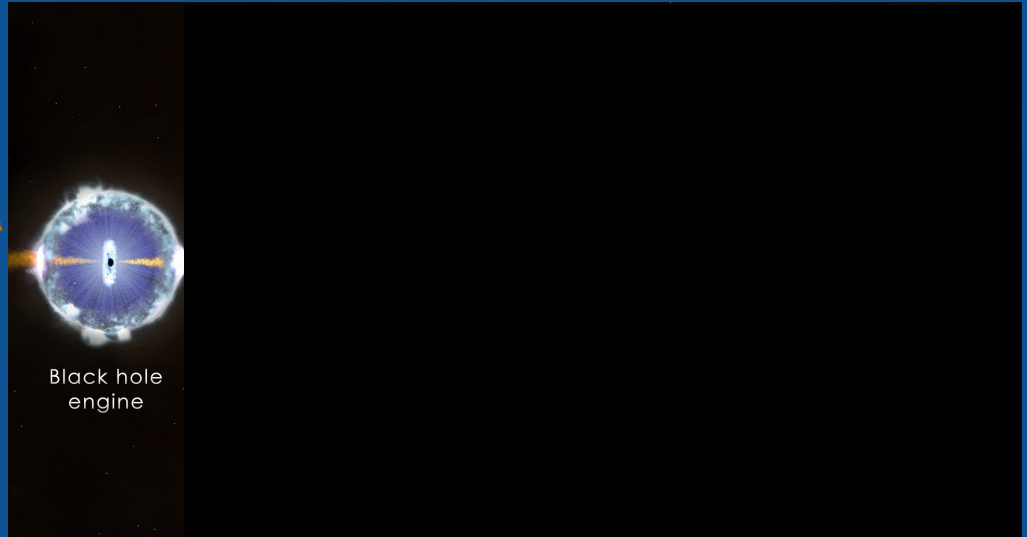
# Context - GRB Phenomena



<https://aasnova.org/>

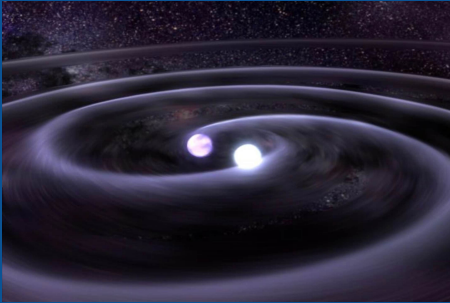


<https://www.nasa.gov/>



NASA/Goddard Space Flight Center

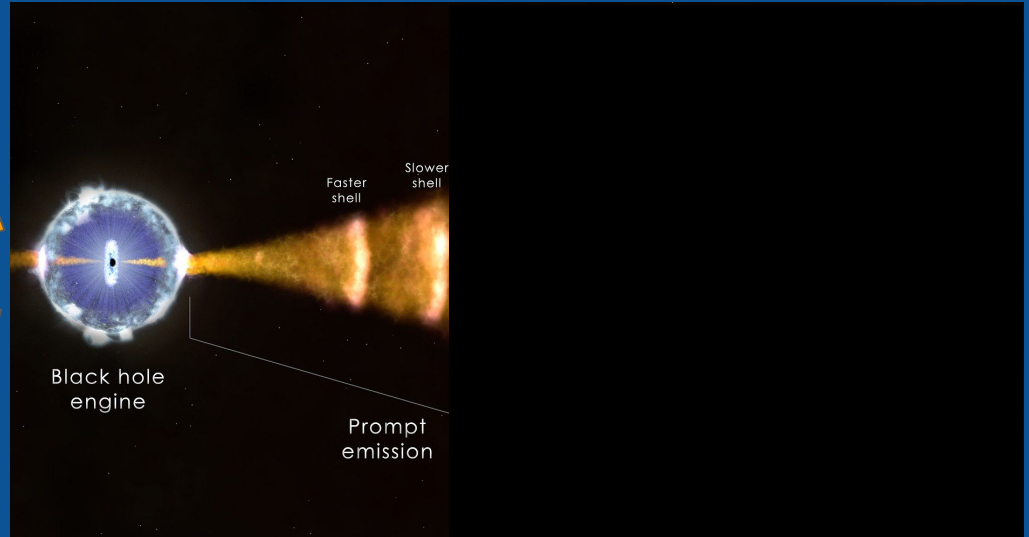
# Context - GRB Phenomena



<https://aasnova.org/>

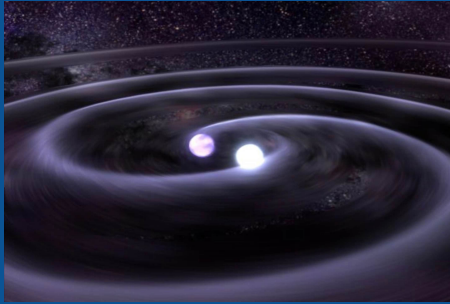


<https://www.nasa.gov/>



NASA/Goddard Space Flight Center

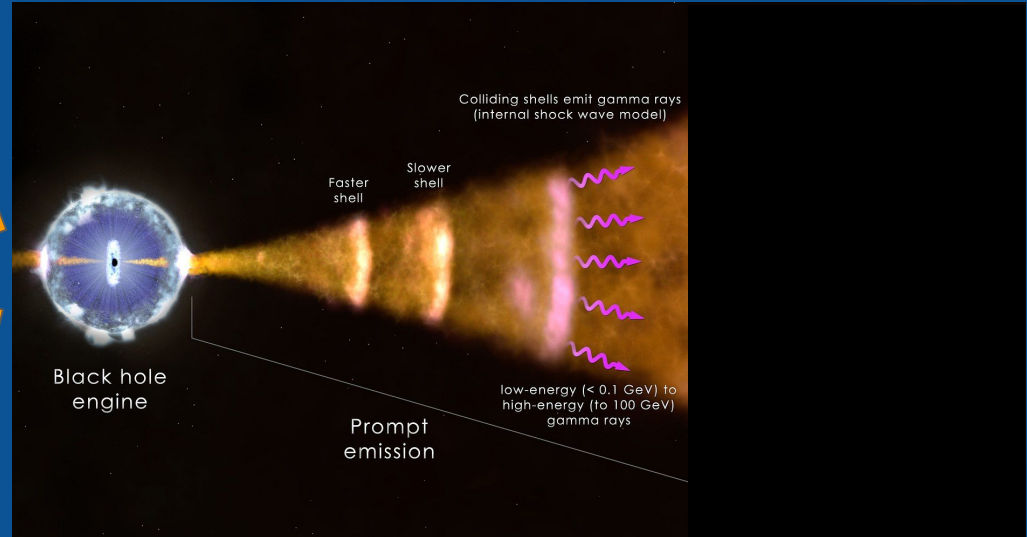
# Context - GRB Phenomena



<https://aasnova.org/>

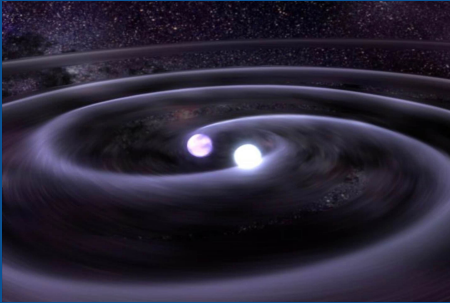


<https://www.nasa.gov/>



NASA/Goddard Space Flight Center

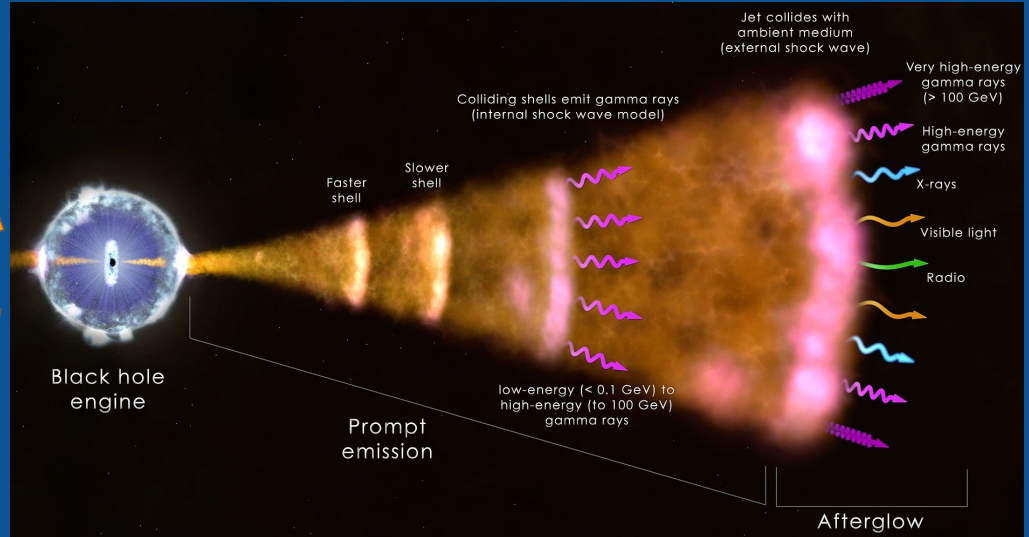
# Context - GRB Phenomena



<https://aasnova.org/>

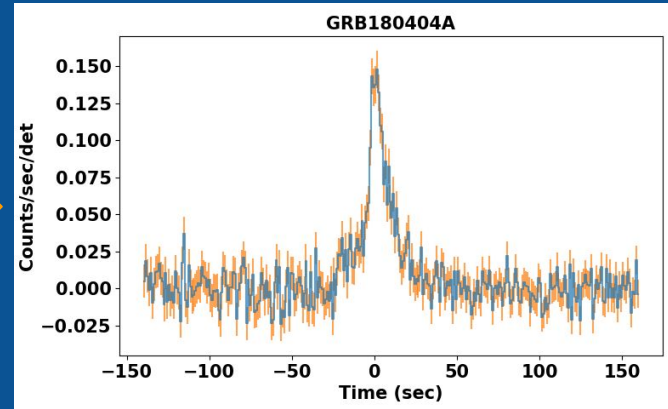
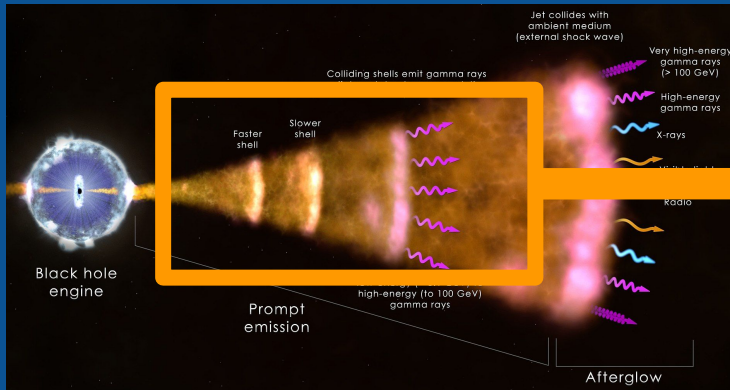


<https://www.nasa.gov/>



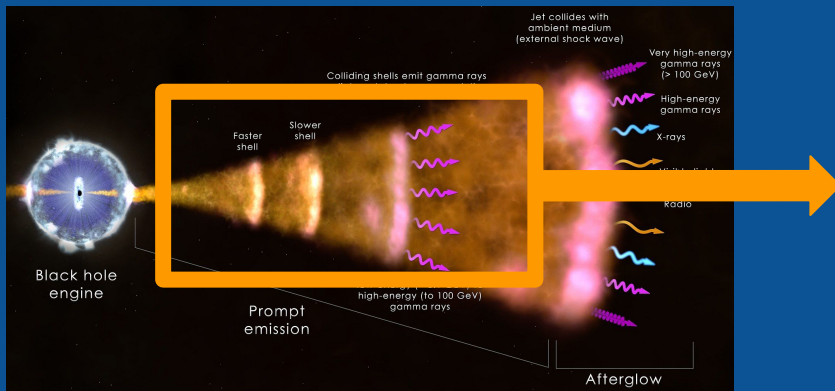
NASA/Goddard Space Flight Center

# Context - GRB Phenomena

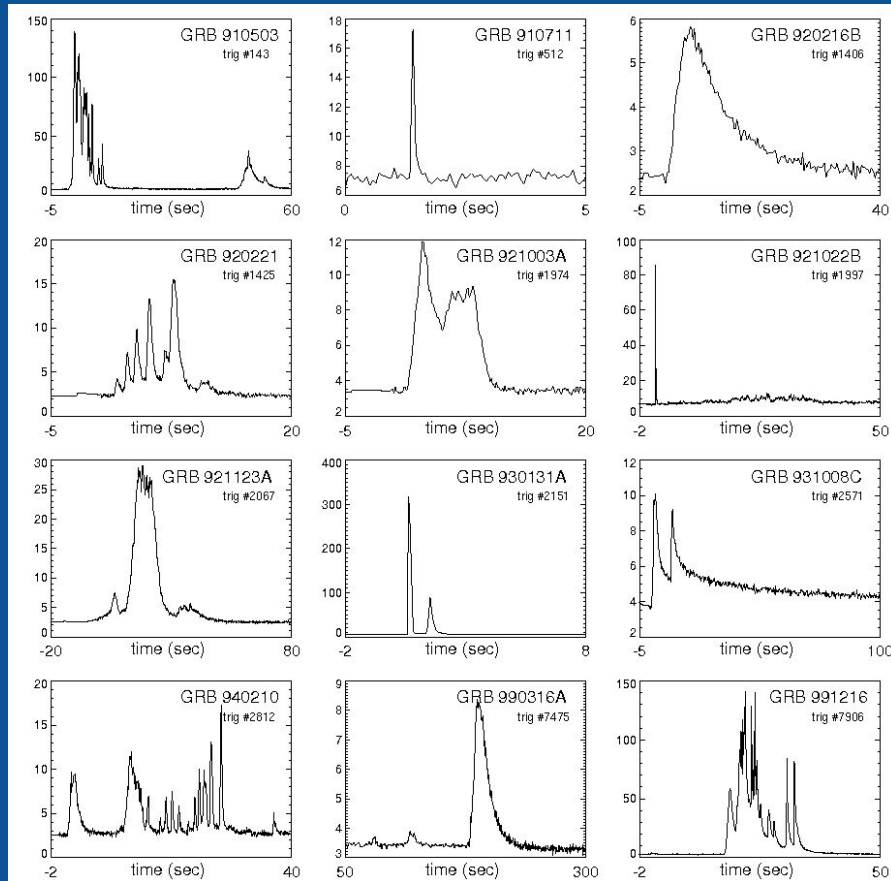


NASA/Goddard Space Flight Center

# Context - GRB Phenomena

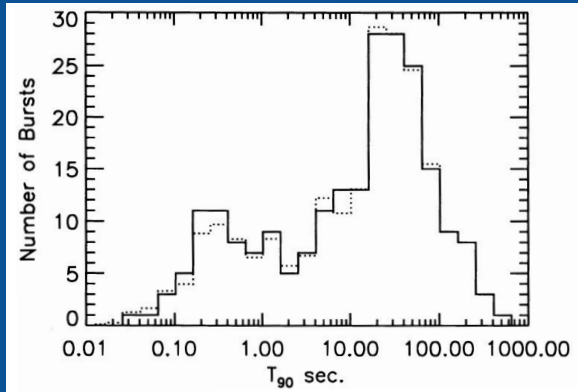


NASA/Goddard Space Flight Center



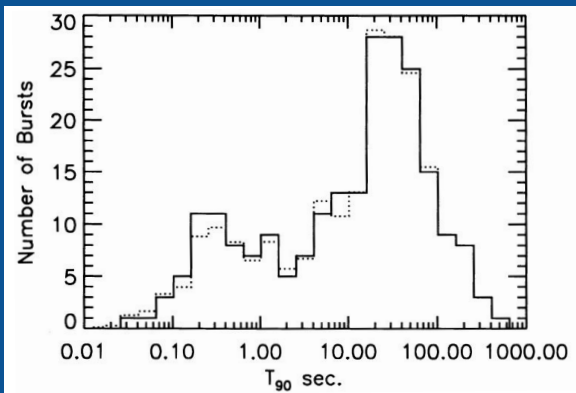


# Instrument Considerations

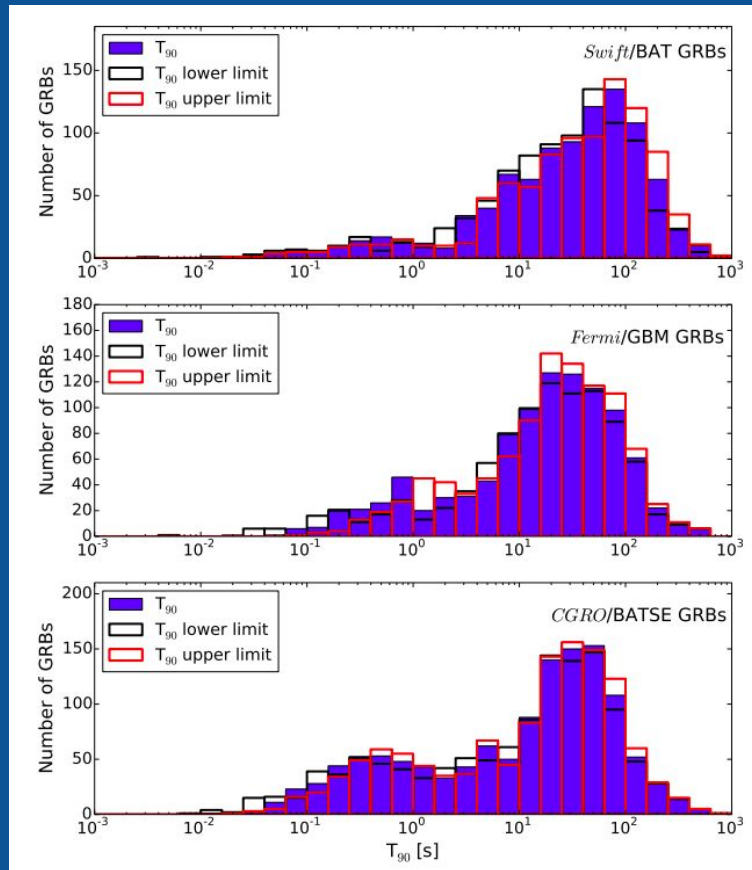


Kouveliotou et al., 1993

# Instrument Considerations

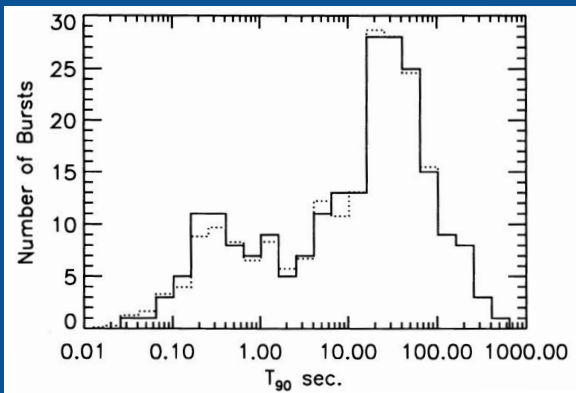


Kouveliotou et al., 1993

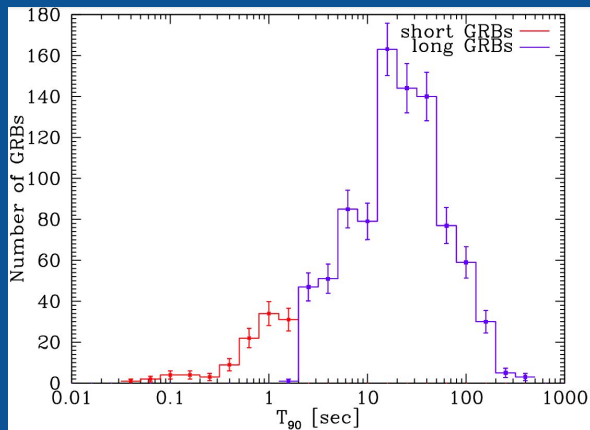


Lien A. et al., 2016

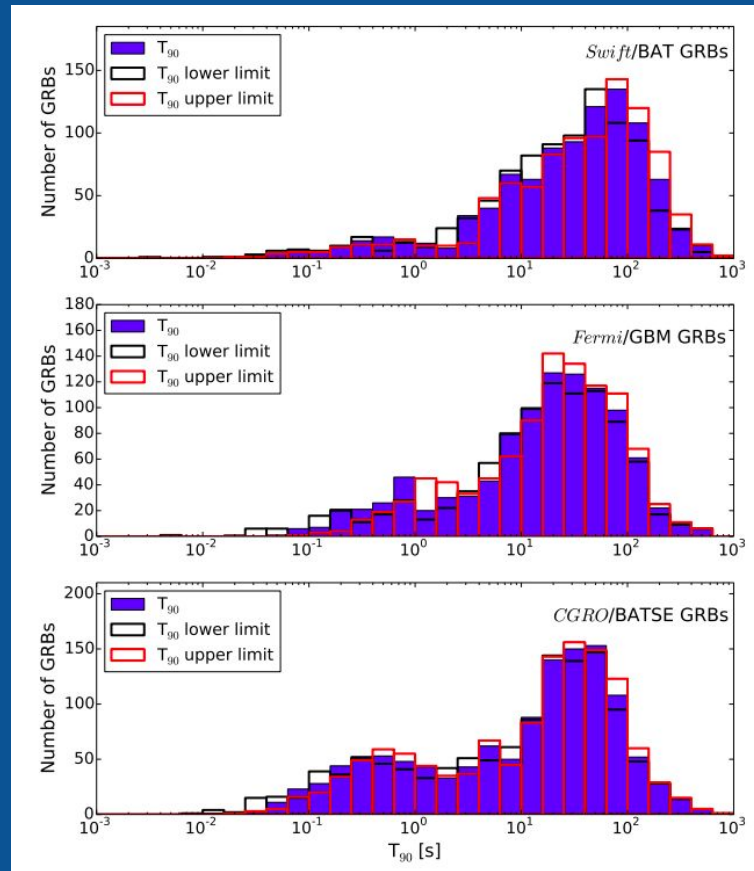
# Instrument Considerations



Kouveliotou et al., 1993



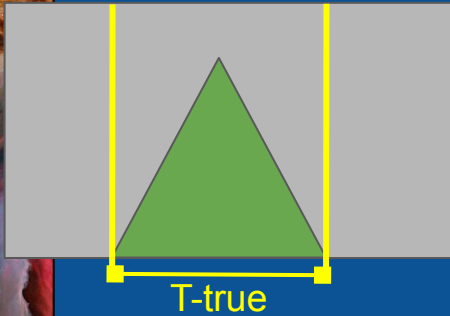
Frontera et al., 2009



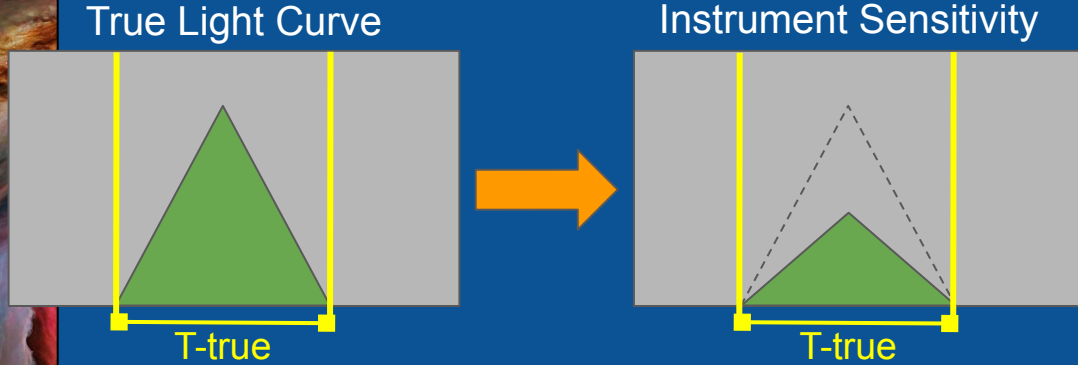
Lien A. et al., 2016

# Instrumental Effects on GRB Light Curves

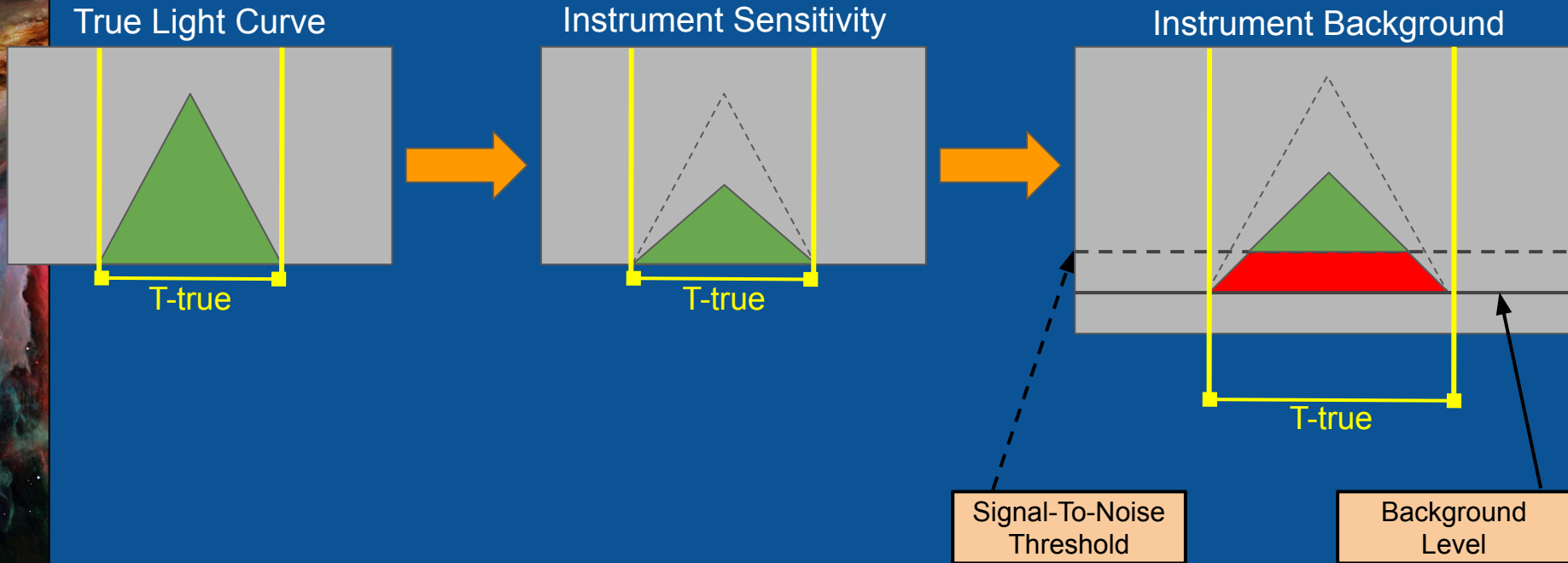
True Light Curve



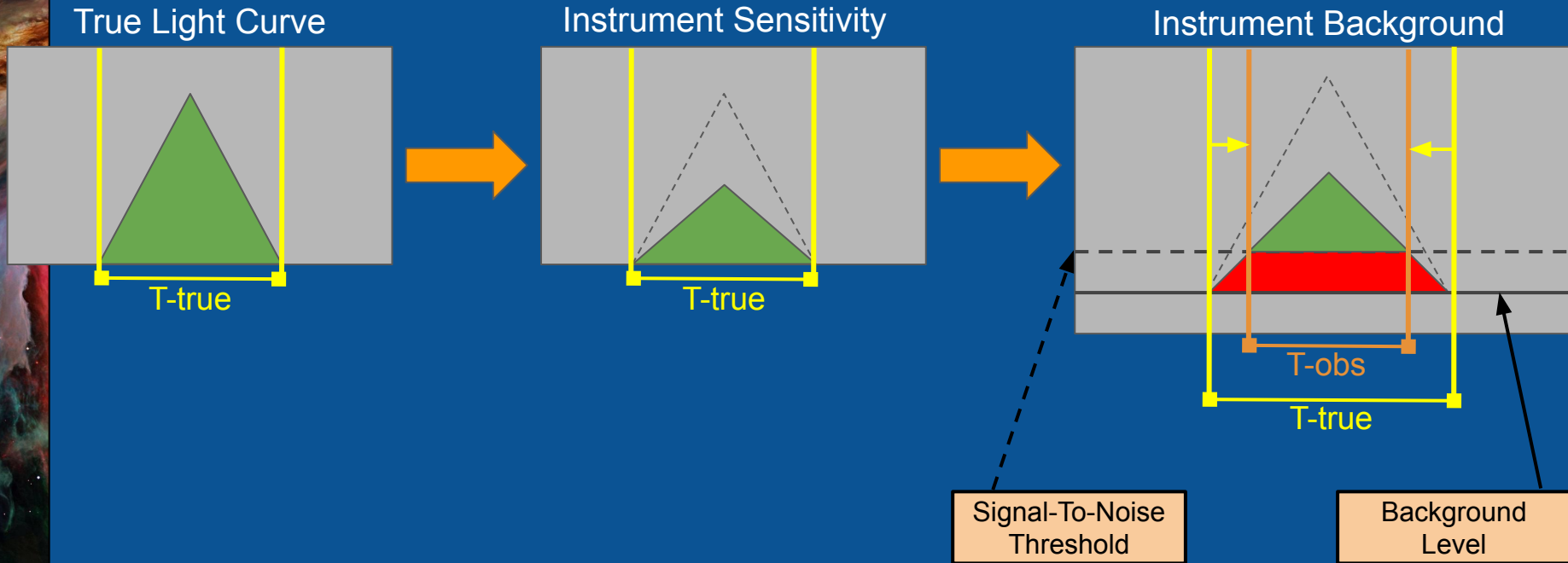
# Instrumental Effects on GRB Light Curves



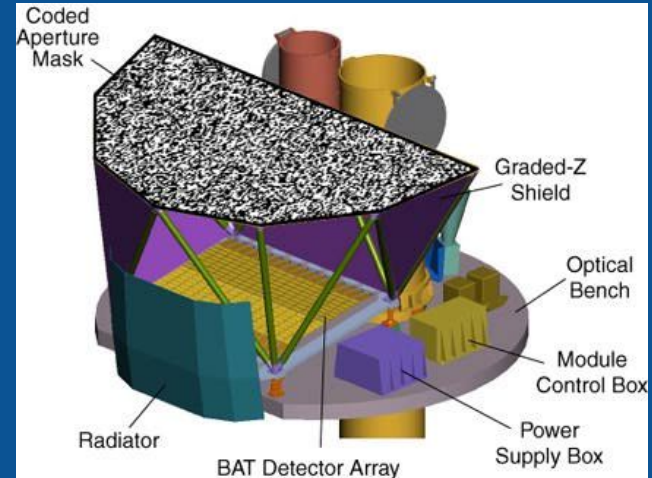
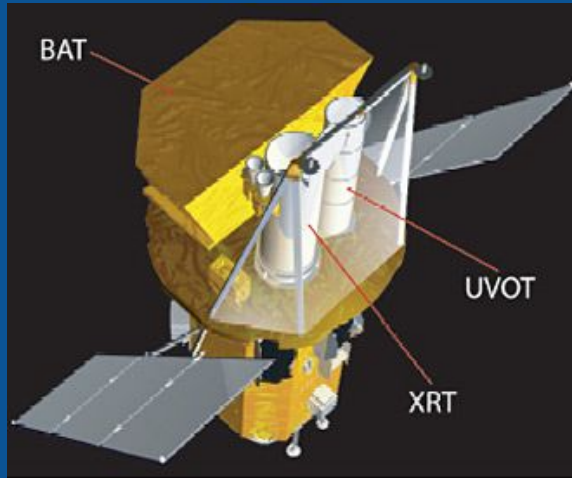
# Instrumental Effects on GRB Light Curves



# Instrumental Effects on GRB Light Curves



# The *Neil Gehrels Swift* Observatory Burst Alert Telescope



[https://www.nasa.gov/mission\\_pages/swift/spacecraft/](https://www.nasa.gov/mission_pages/swift/spacecraft/)

[https://swift.gsfc.nasa.gov/about\\_swift/bat\\_desc.html](https://swift.gsfc.nasa.gov/about_swift/bat_desc.html)



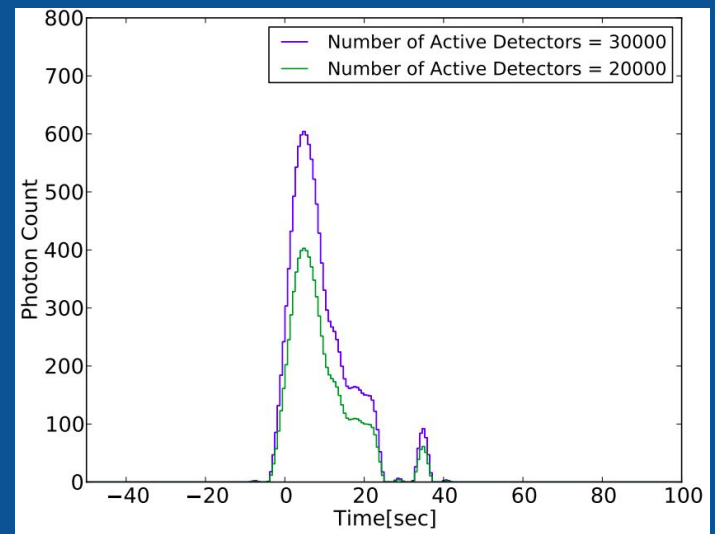
# Swift/BAT Instrumental Parameters

Relevant Parameters:

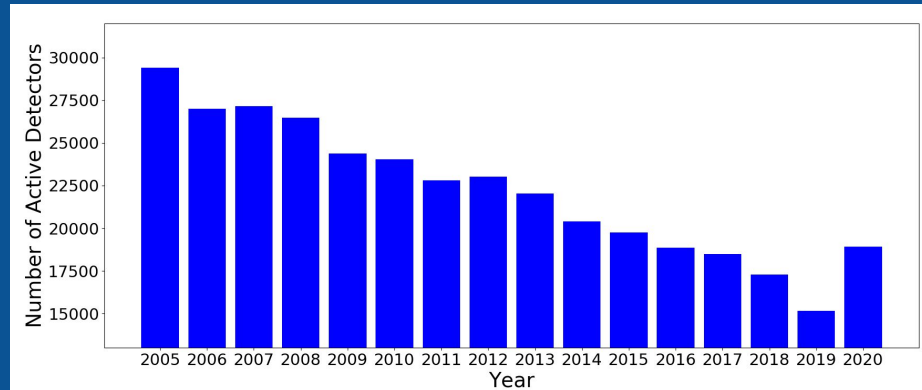
- **Number of Active Detectors (NDETS)**
- Incident angle (PCODE)
- Background

Not Relevant:

- Energy band



Lien A. et al., 2014



# Swift/BAT Instrumental Parameters

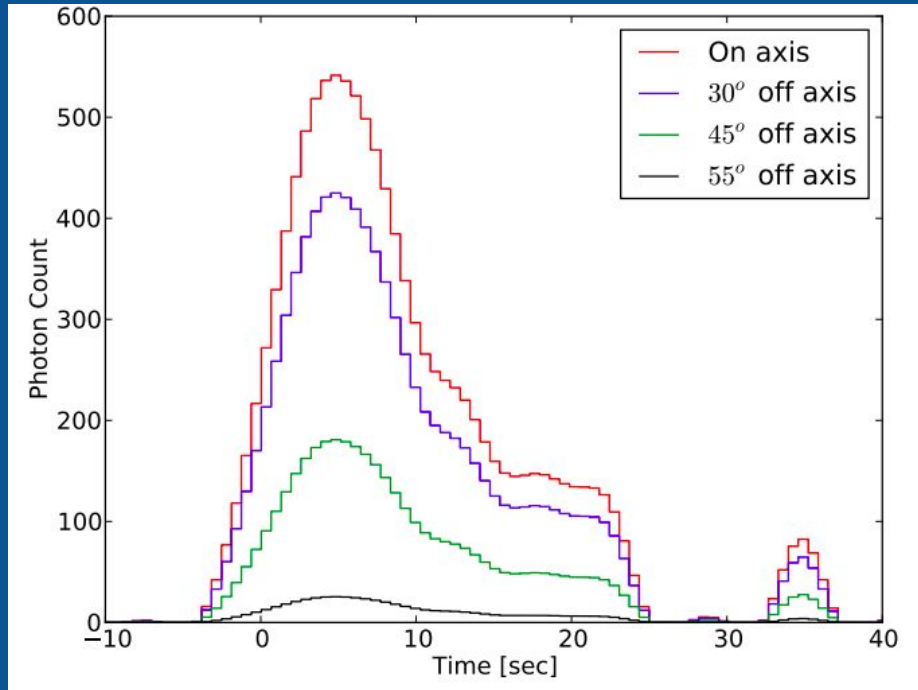
Relevant Parameters:

- Number of Active Detectors (NDETS)
- **Incident angle (PCODE)**

- Background

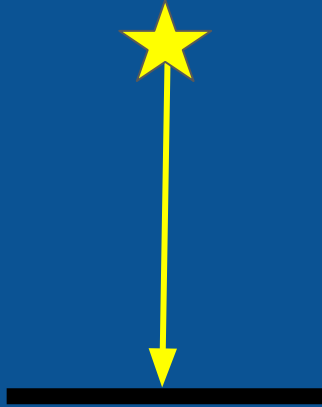
Not Relevant:

- Energy band

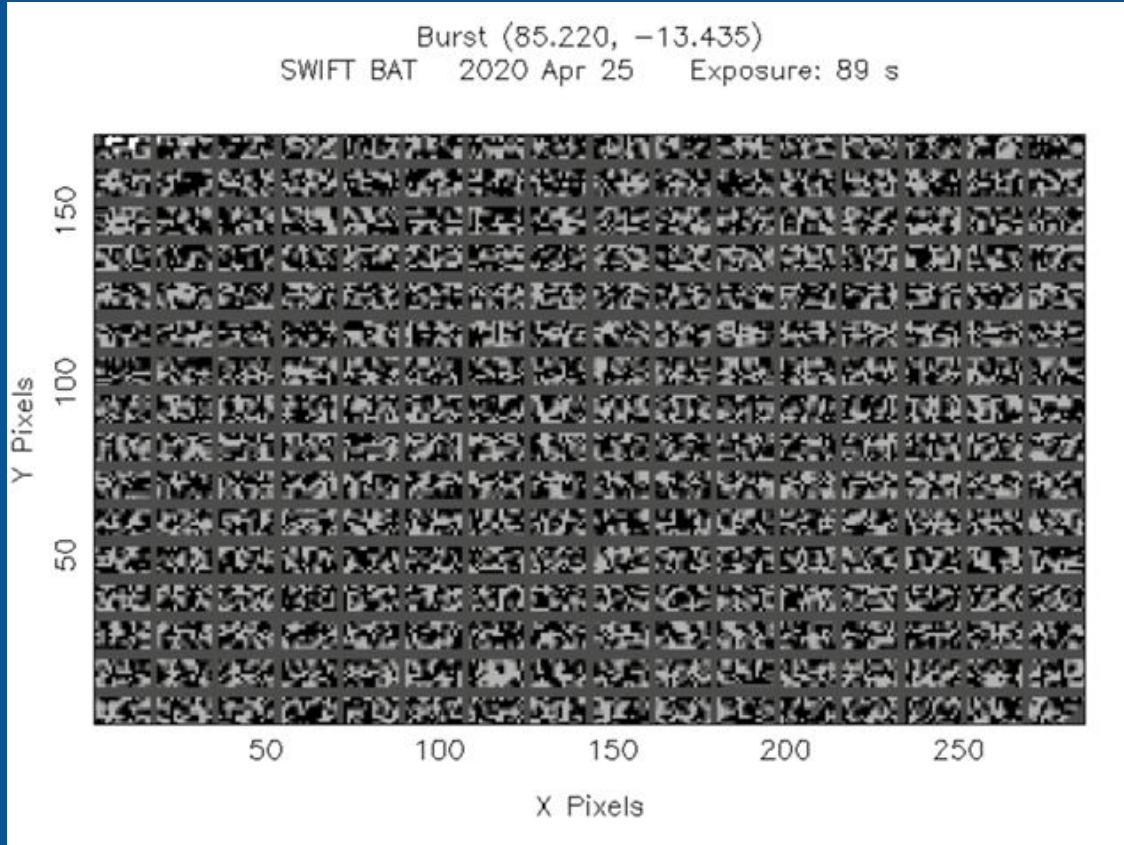


Lien A. et al., 2014

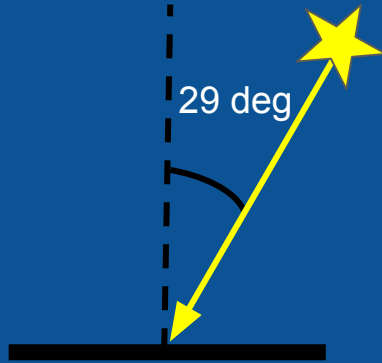
# PCODE and Incident Angle



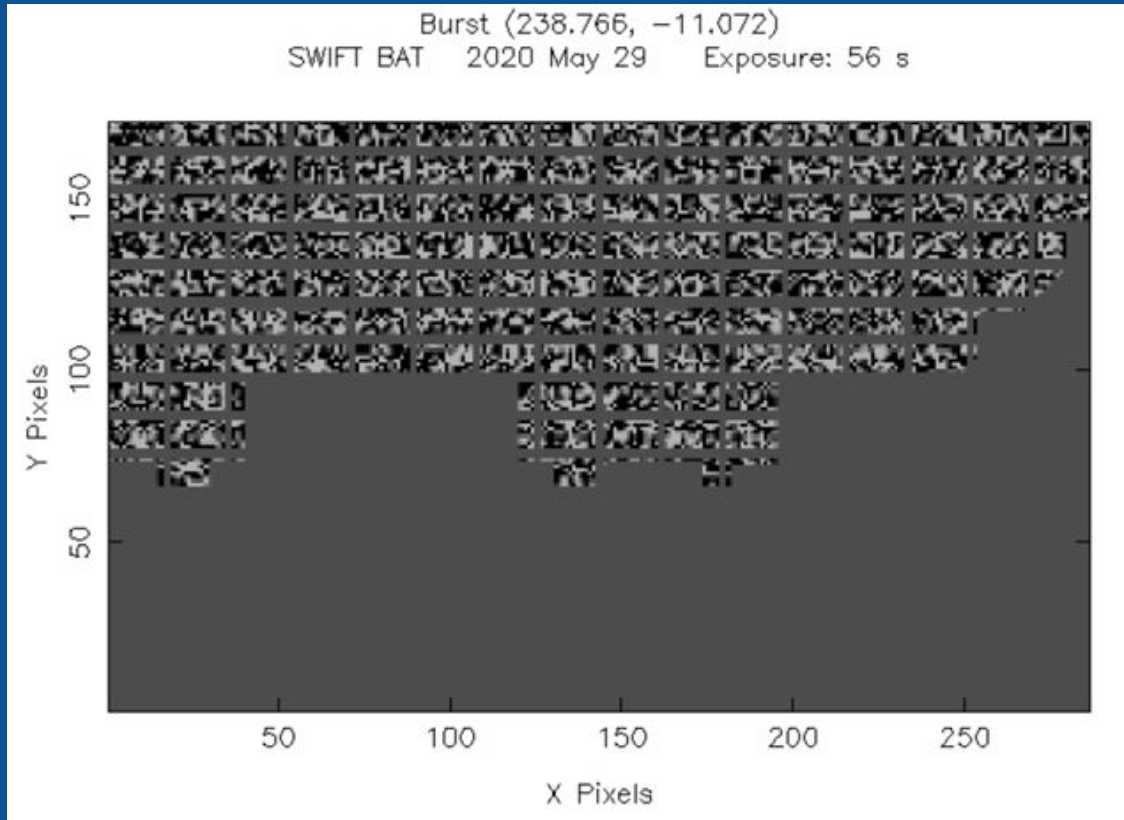
Incident Angle = 0 deg  
PCODE = 1



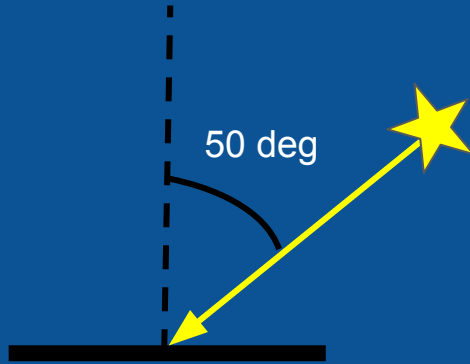
# PCODE and Incident Angle



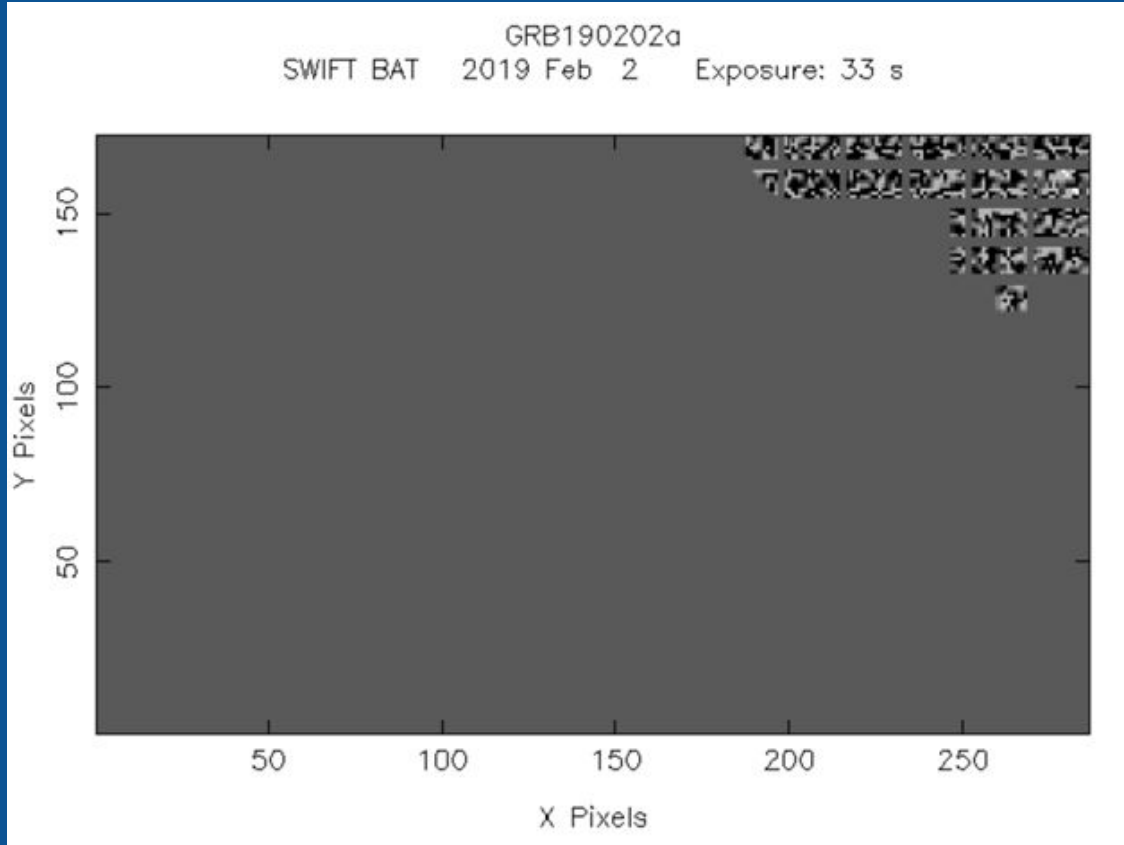
Incident Angle = 29 deg  
PCODE = 0.5



# PCODE and Incident Angle

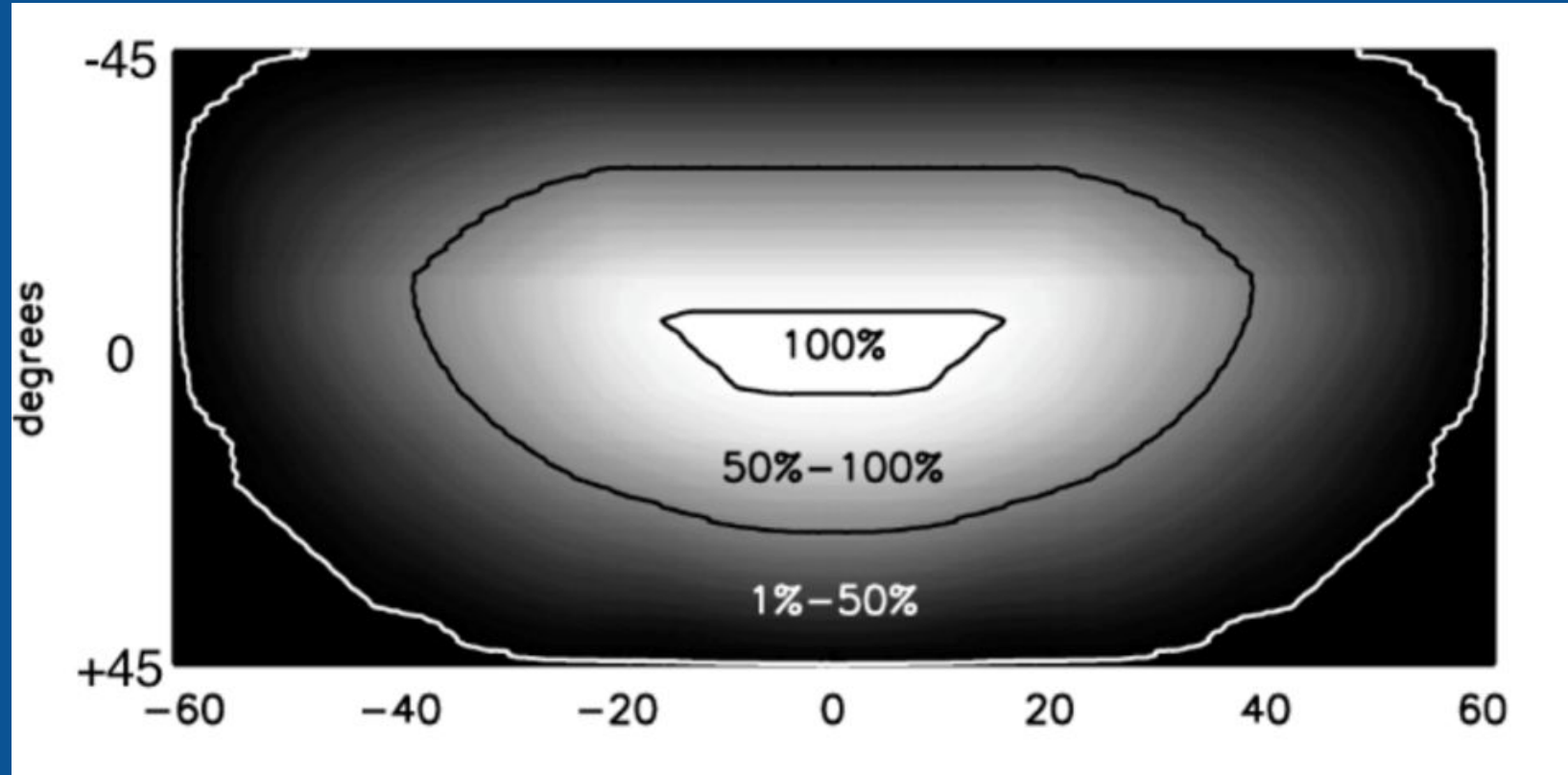


Incident Angle = 50 deg  
PCODE = 0.04



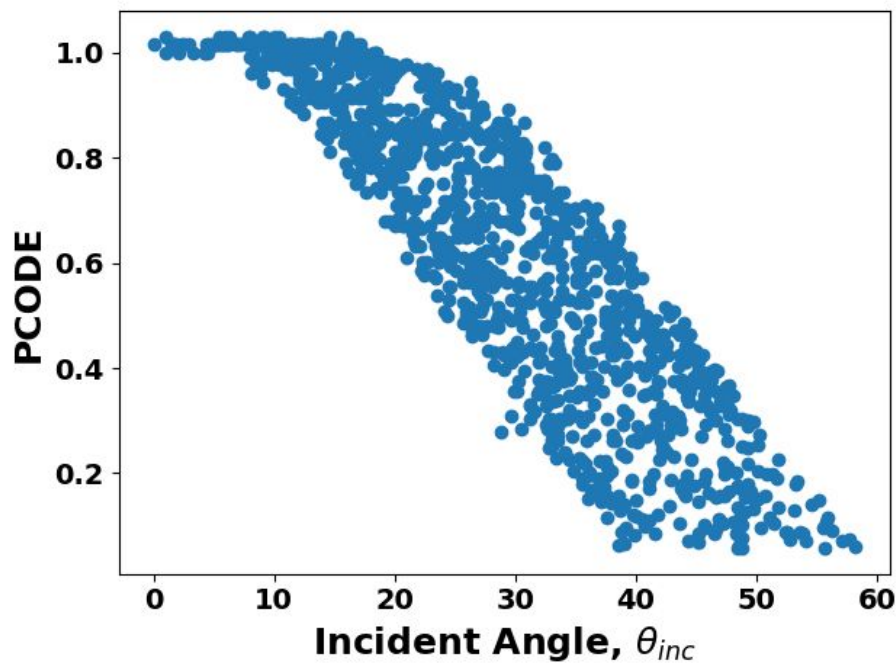
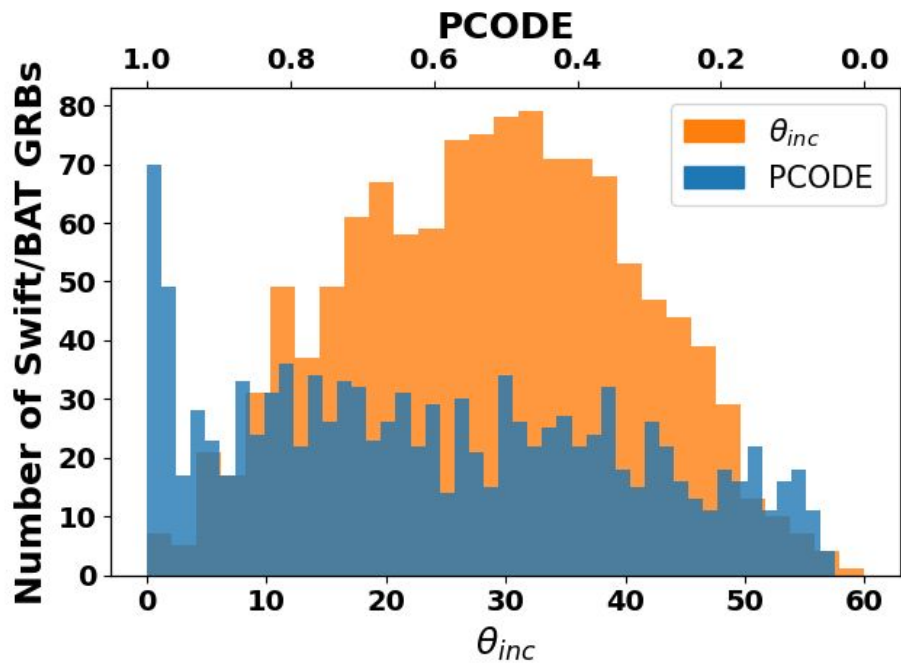
# PCODE vs BAT Field of View

Not exactly one-to-one with incident angle



# PCODE vs BAT Field of View

Not exactly one-to-one with incident angle



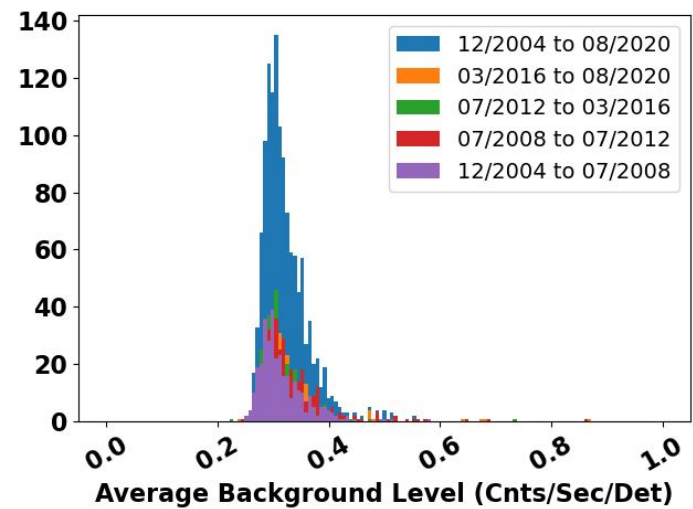
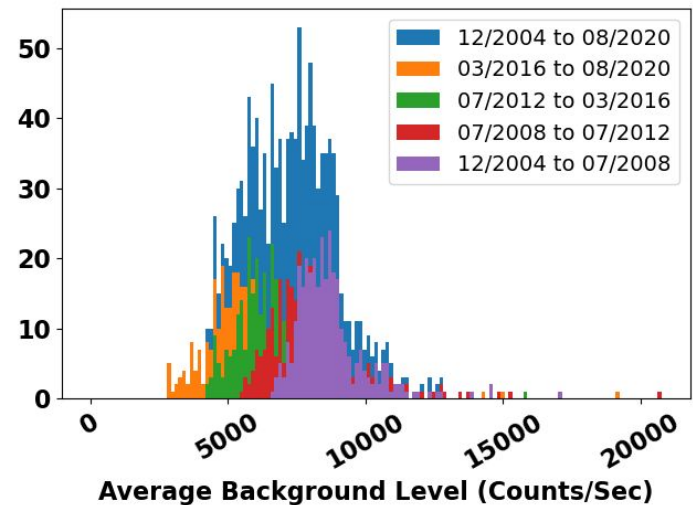
# Swift/BAT Instrumental Parameters

Relevant Parameters:

- Number of Active Detectors (NDETS)
- Incident angle (PCODE)
- **Background**

Not Relevant:

- Energy band





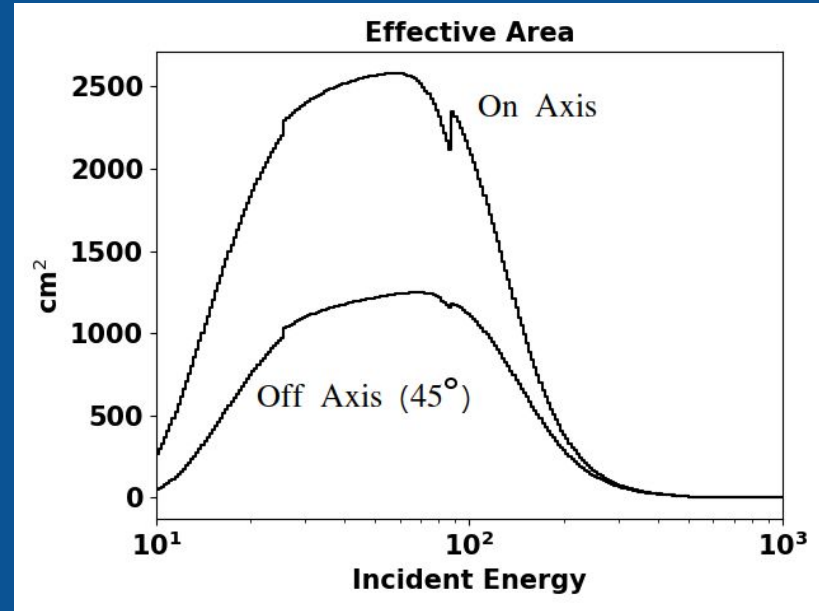
# Swift/BAT Instrumental Parameters

Relevant Parameters:

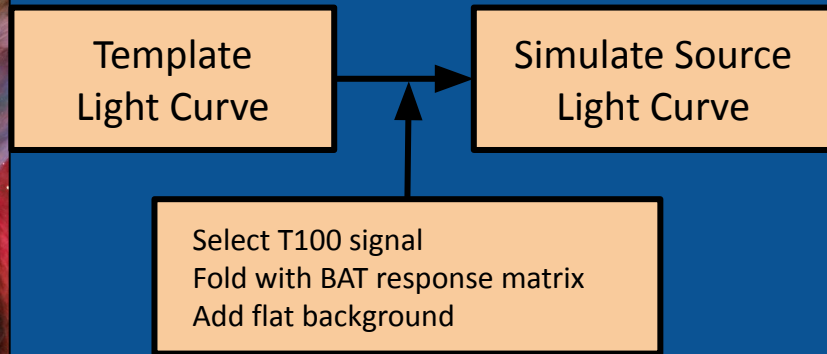
- Number of Active Detectors (NDETS)
- Incident angle (PCODE)
- Background

Not Relevant:

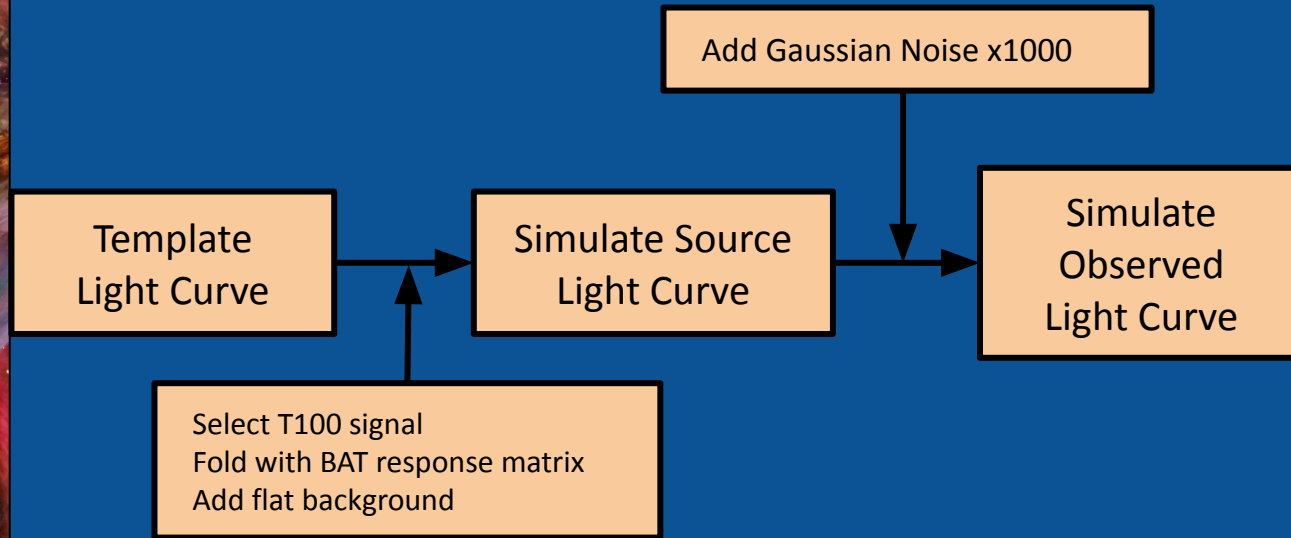
- **Energy band**



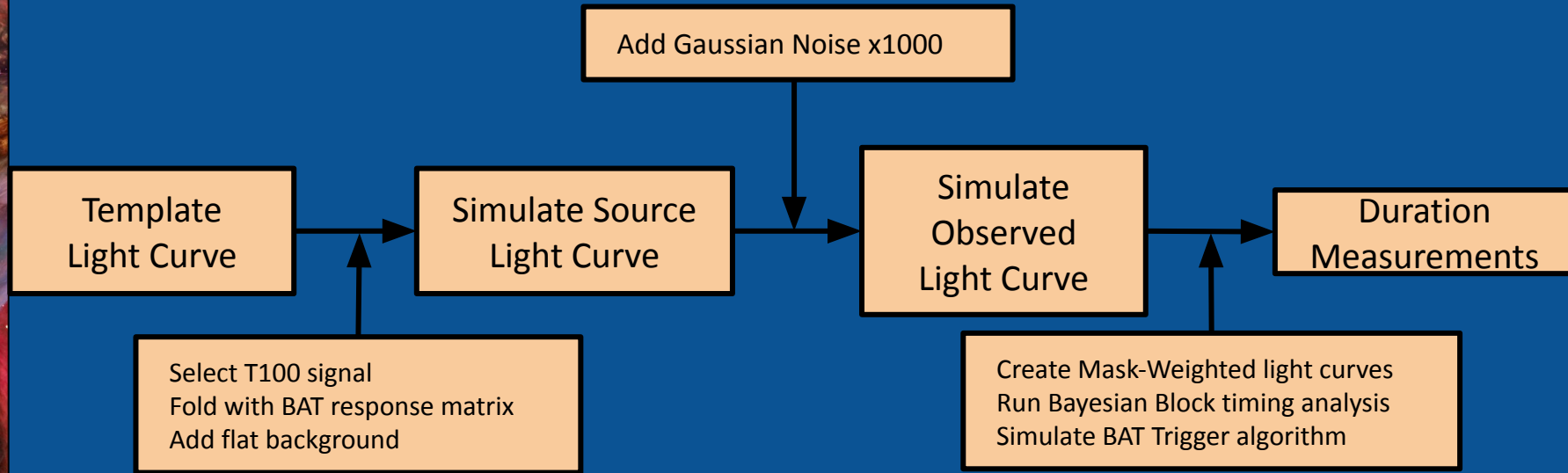
# Simulation Method Summary



# Simulation Method Summary



# Simulation Method Summary



# Light Curve Sample for Simulations

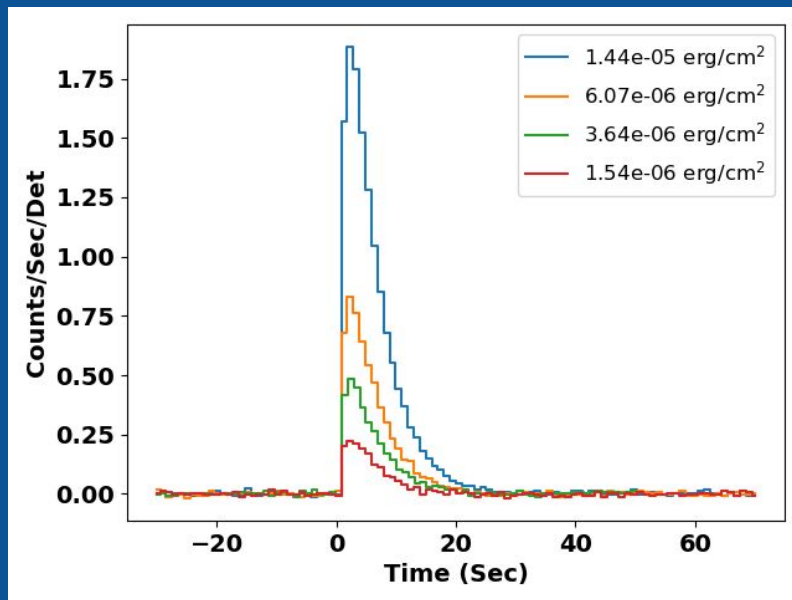
## FRED Light Curves

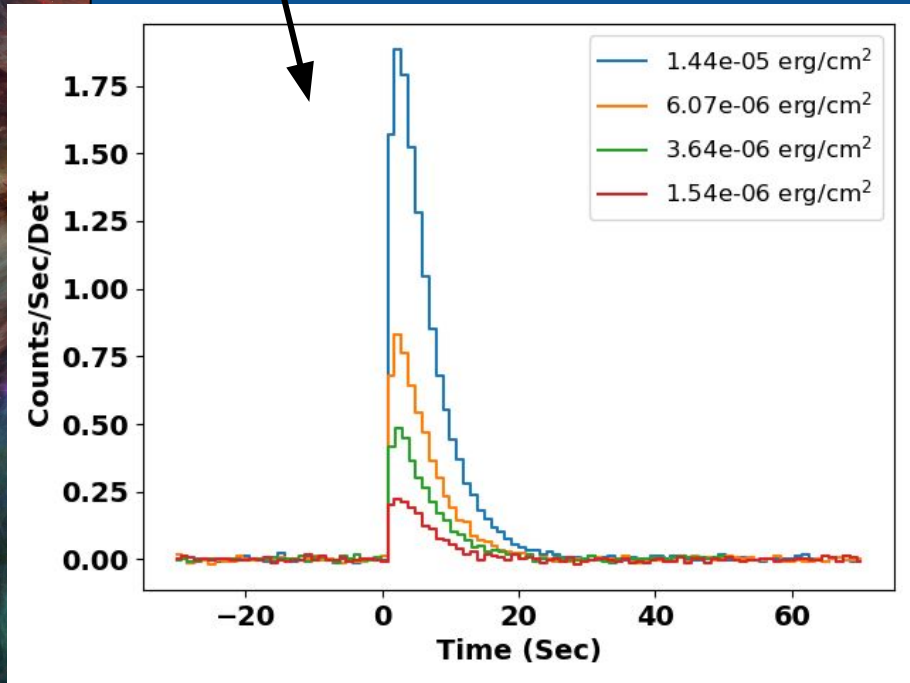
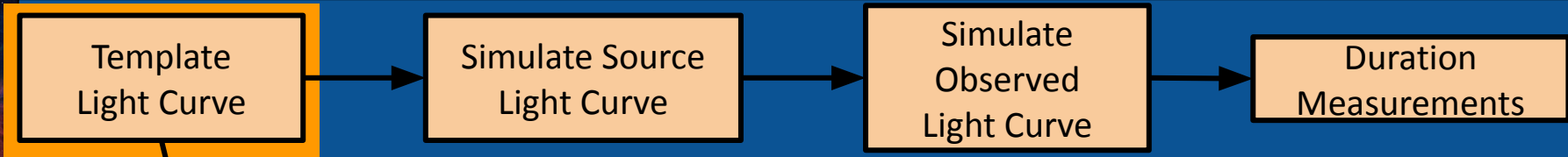
$$I(t) = A\lambda e^{[-\tau_1/(t-t_s) - (t-t_s)/\tau_2]}$$

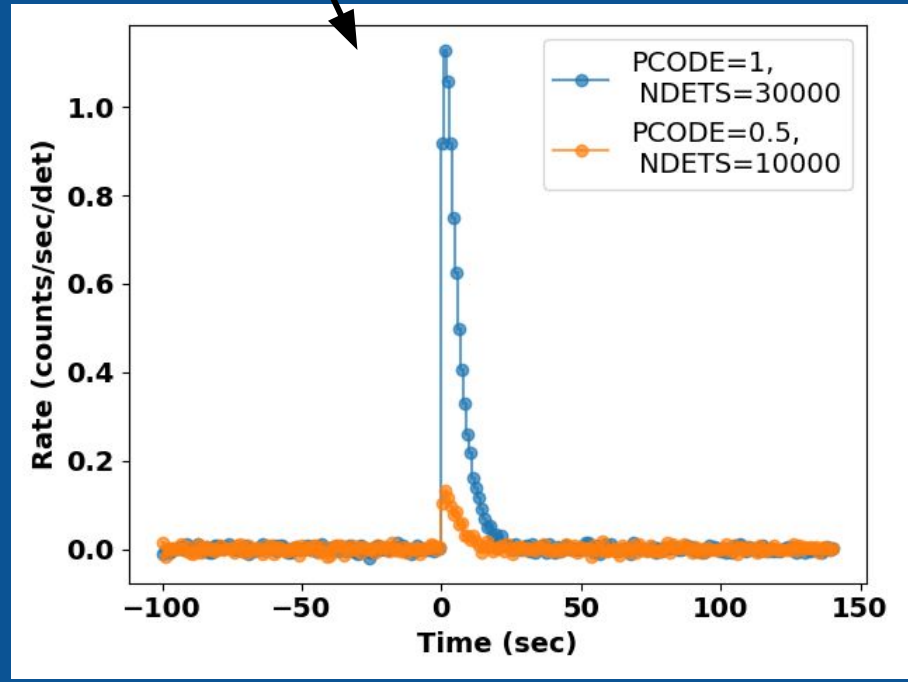
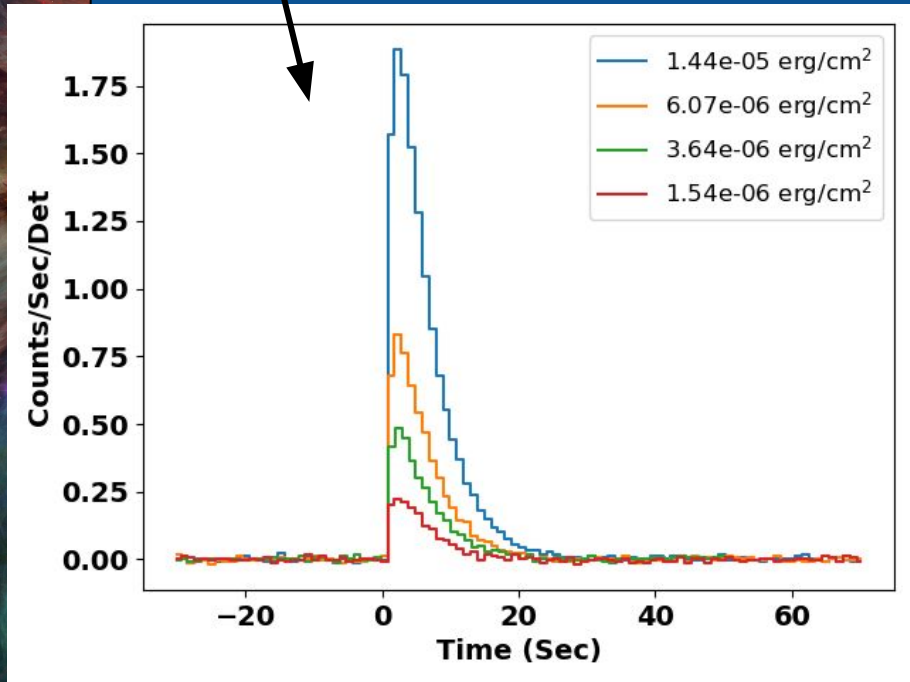
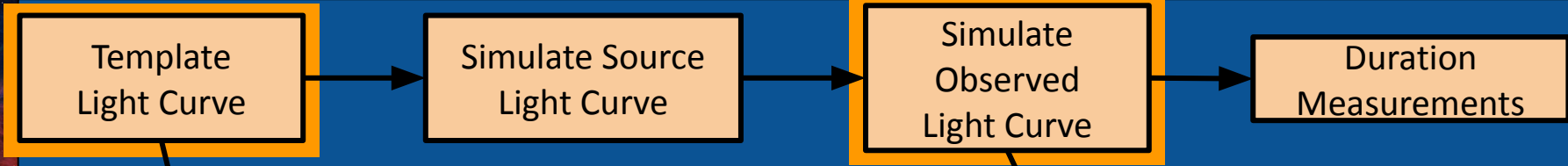
Hakkila and Preece, 2014

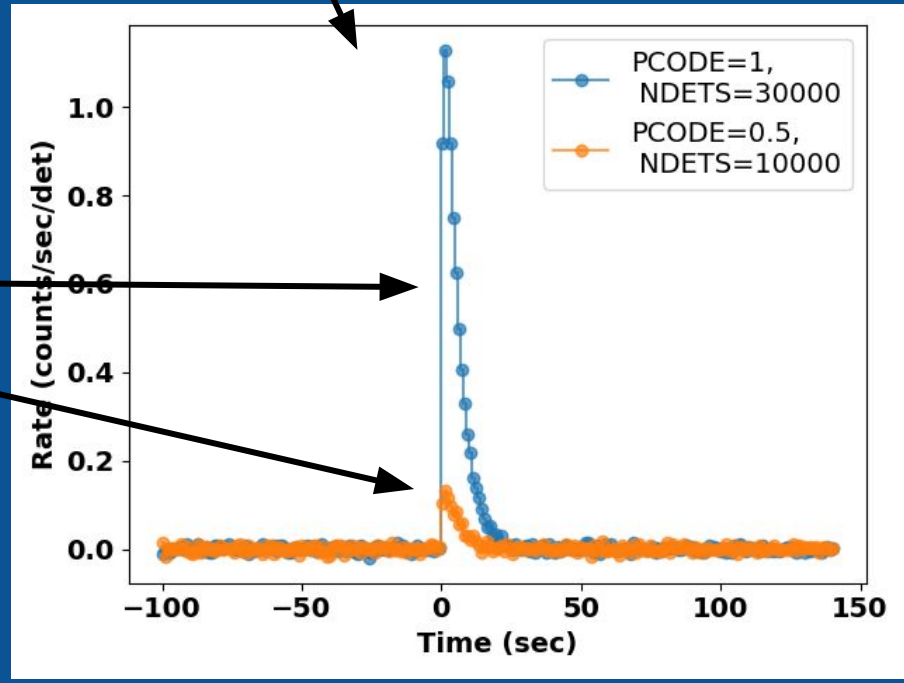
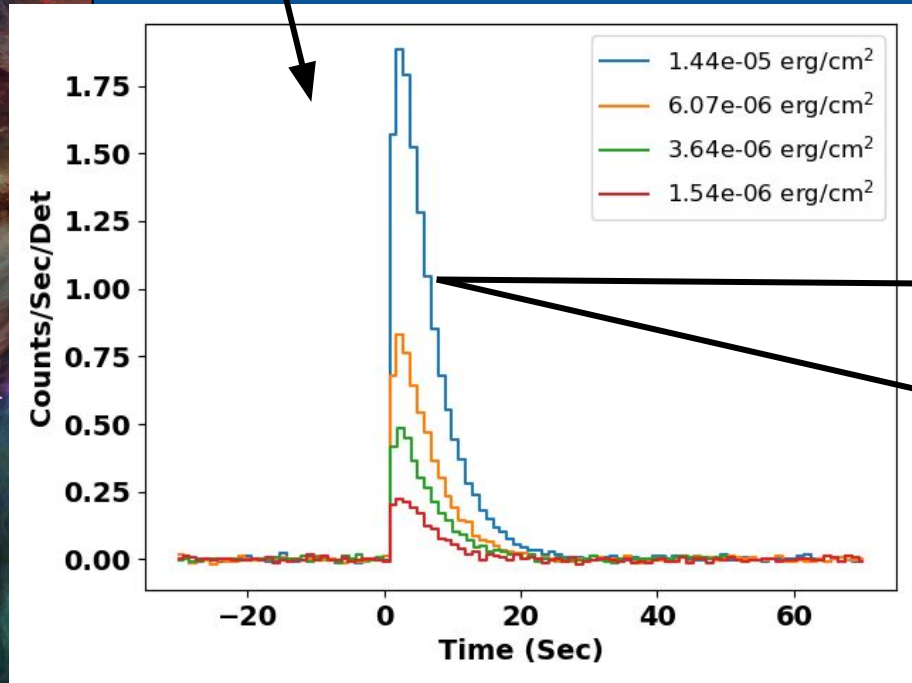
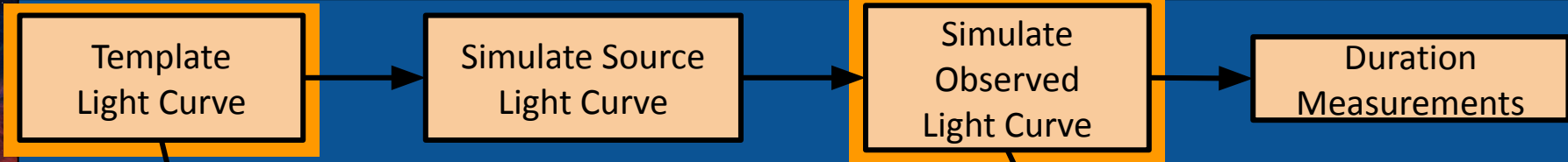
## Observed GRB Light Curves:

GRB050219A  
GRB051111  
GRB071010B  
GRB090510  
GRB110422A  
GRB120119A  
GRB150314A  
GRB160314A









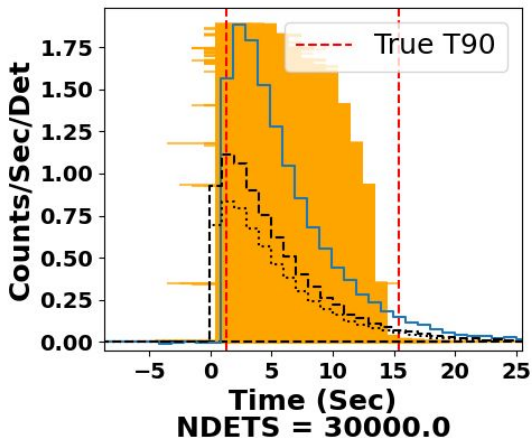


Template  
Light Curve

Simulate Source  
Light Curve

Simulate  
Observed  
Light Curve

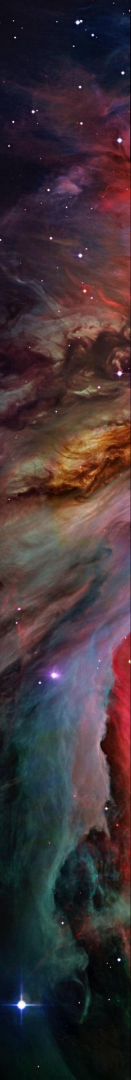
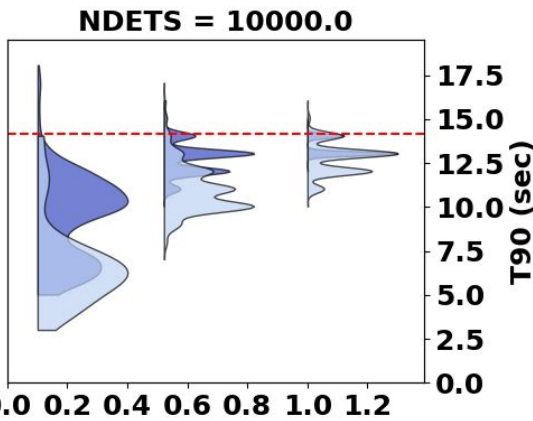
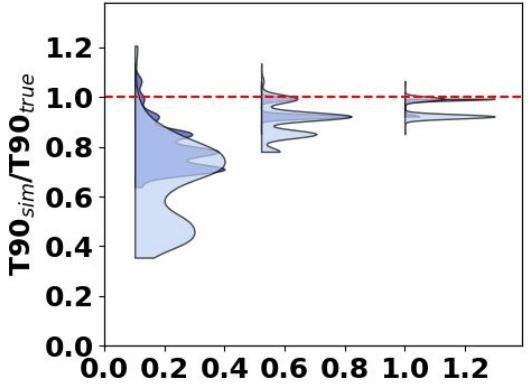
Duration  
Measurements



Ideal FRED Light Curve

Background (c/s):

- 1000.0
- 9500.0

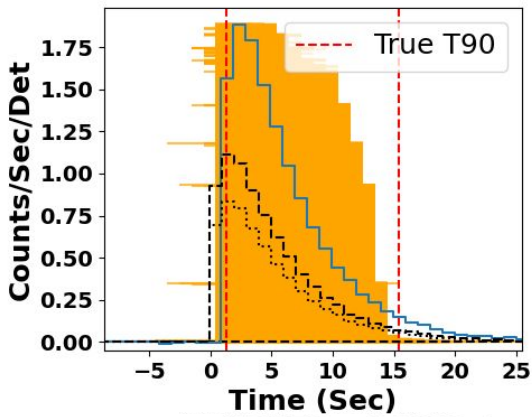


Template  
Light Curve

Simulate Source  
Light Curve

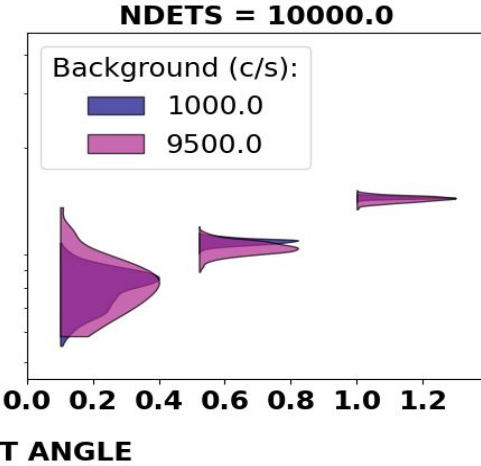
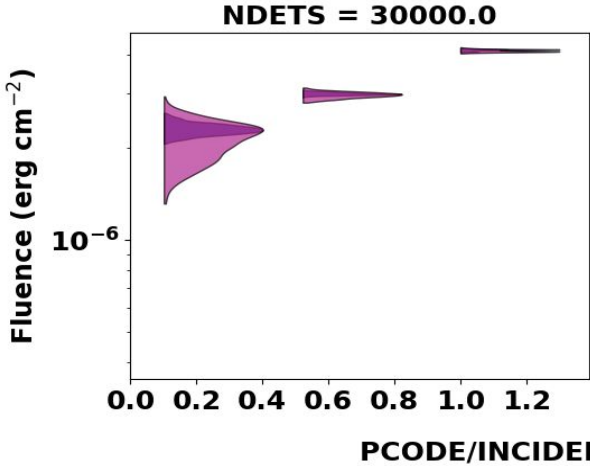
Simulate  
Observed  
Light Curve

Duration  
Measurements

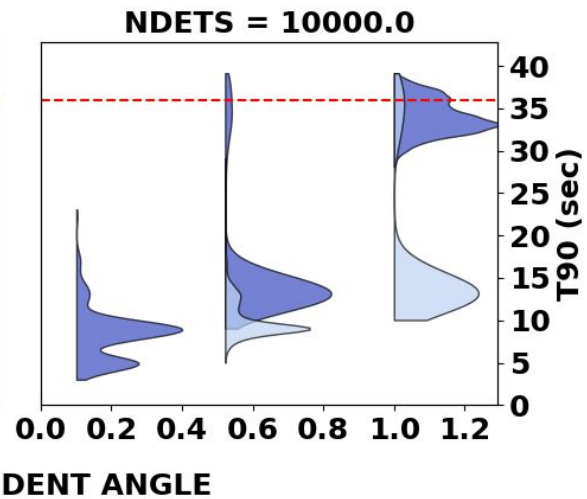
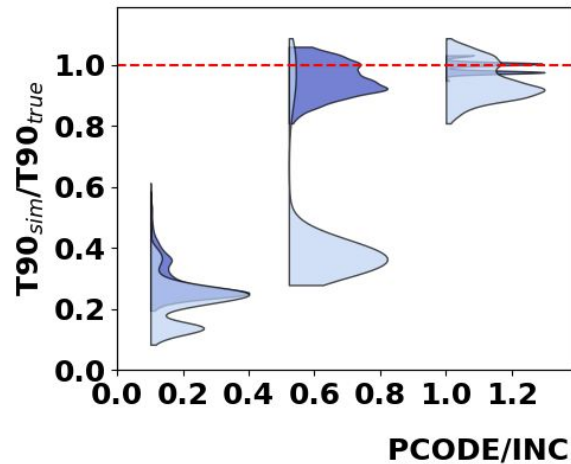
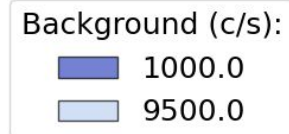
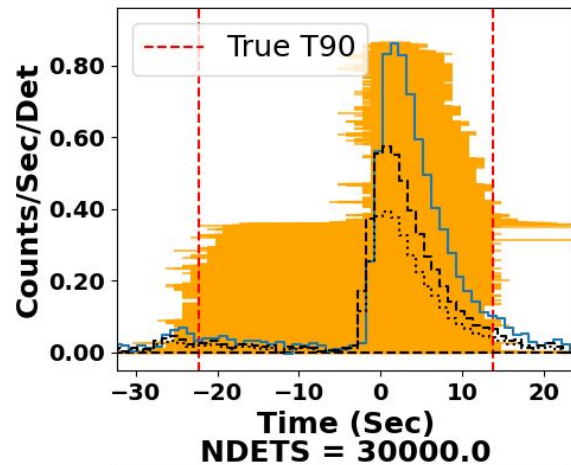


Ideal FRED Light Curve

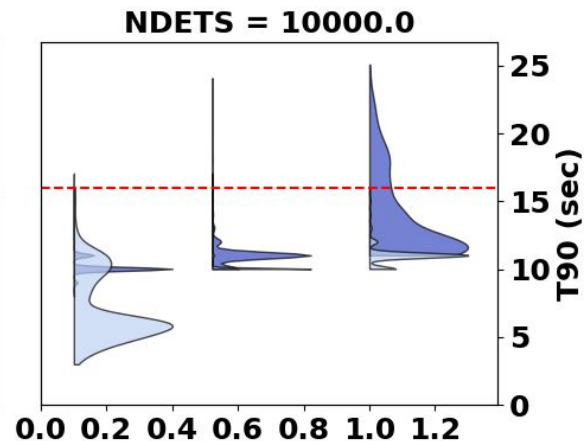
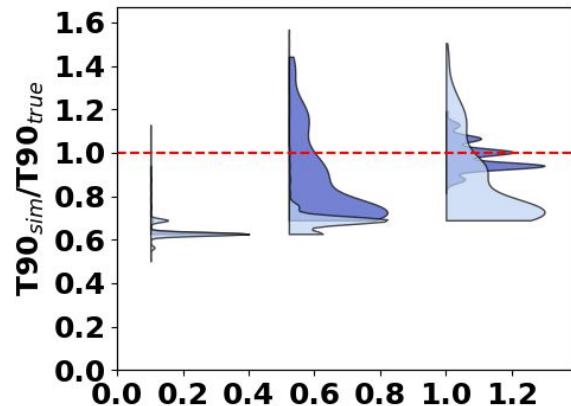
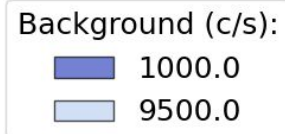
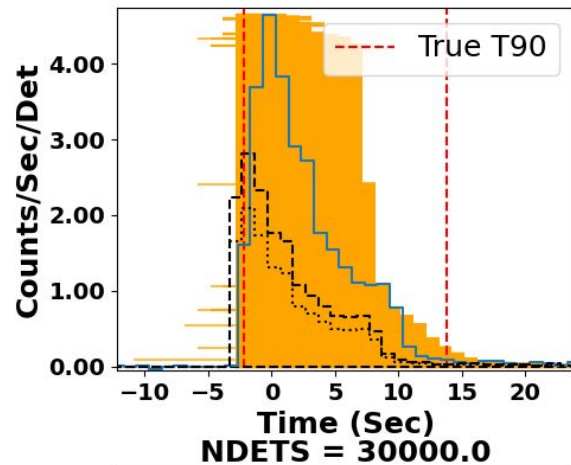
Background (c/s):  
1000.0  
9500.0



# GRB071010B: Early, Dim Emission

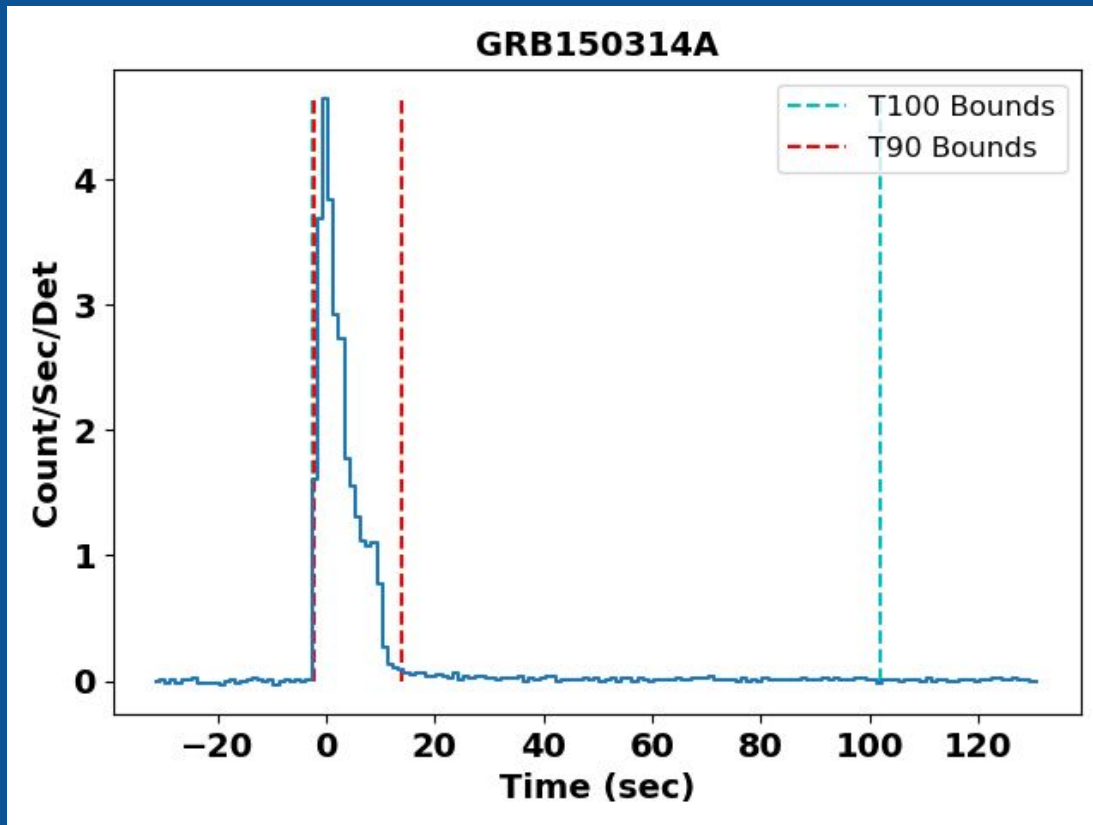


# GRB150314A: LGRB with Extended Emission

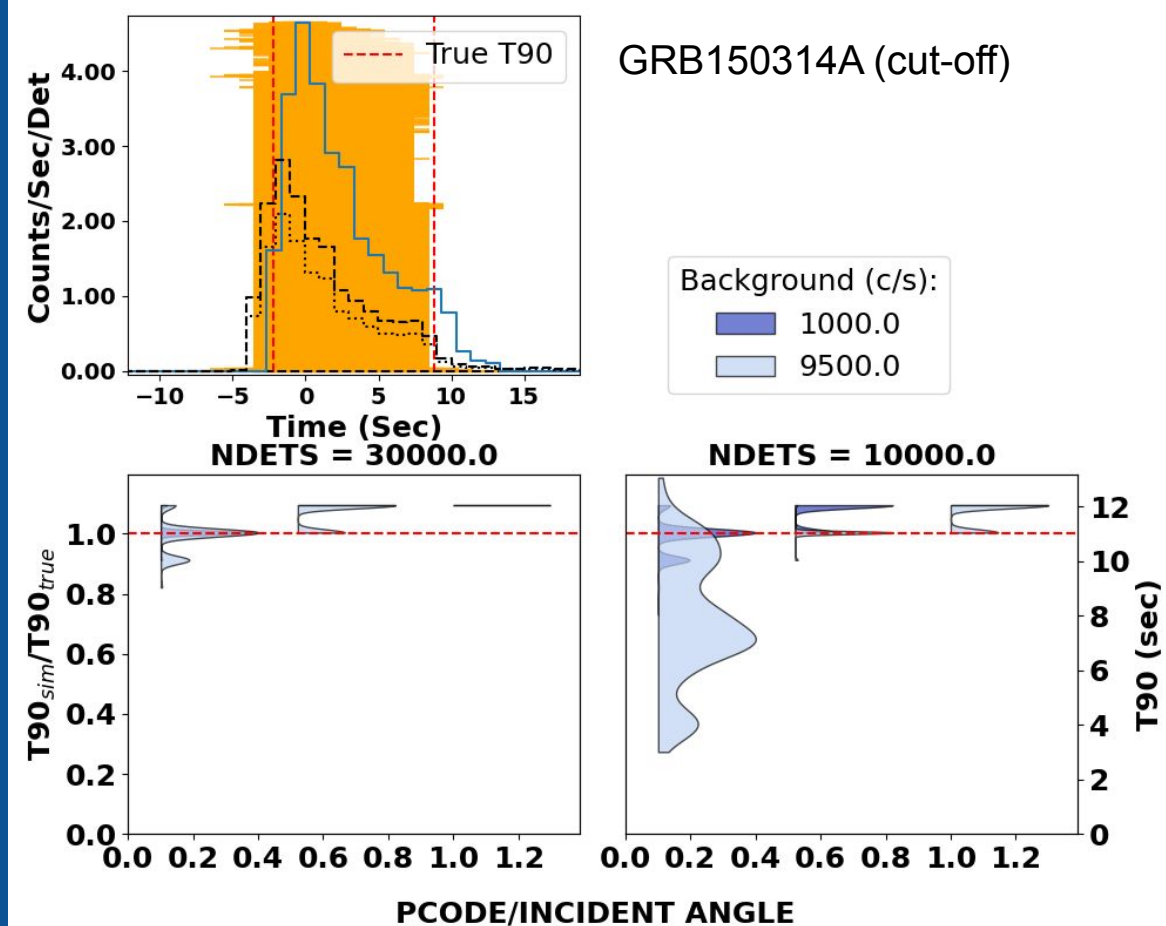
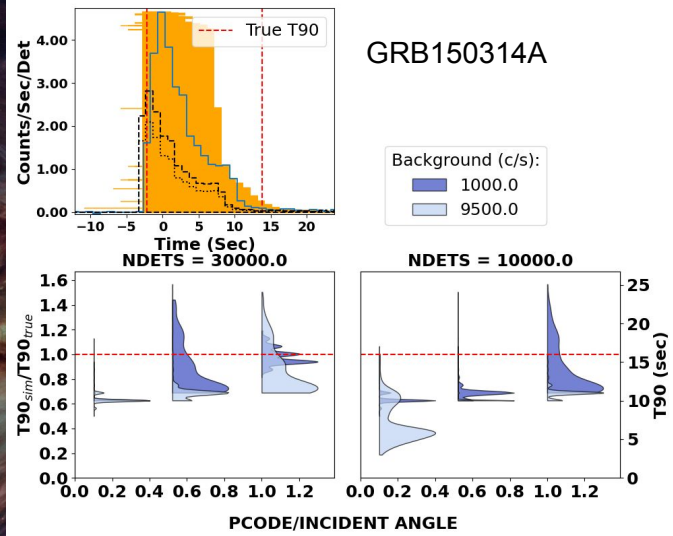


PCODE/INCIDENT ANGLE

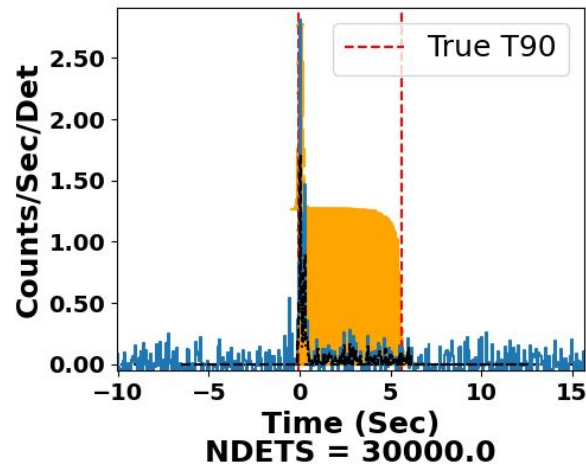
# GRB150314A: LGRB with Extended Emission



# GRB150314A: LGRB with Extended Emission



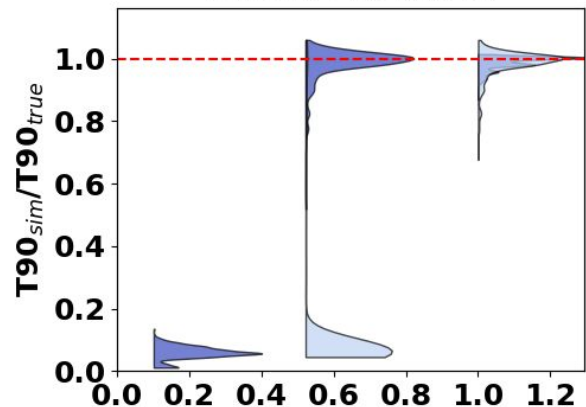
# GRB090510: SGRB with trailing Emission



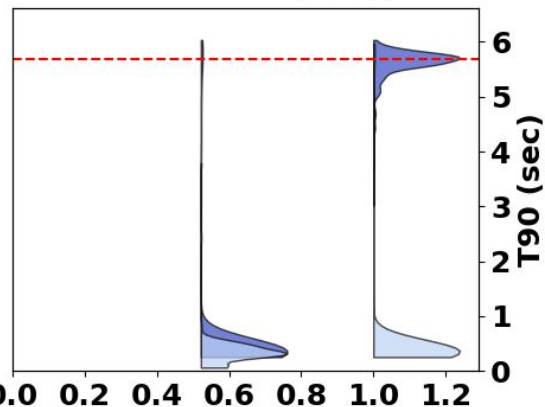
Background (c/s):

1000.0

9500.0

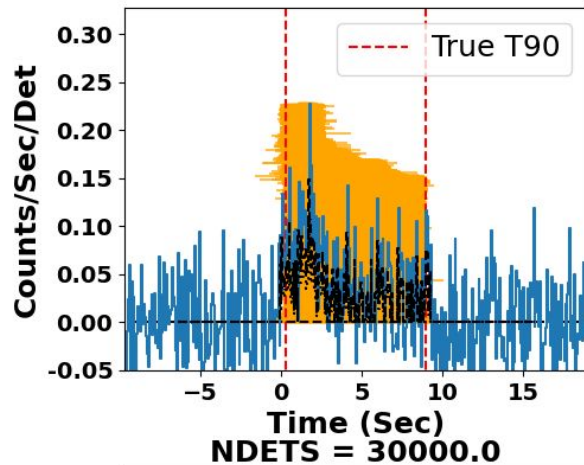


NDETS = 10000.0



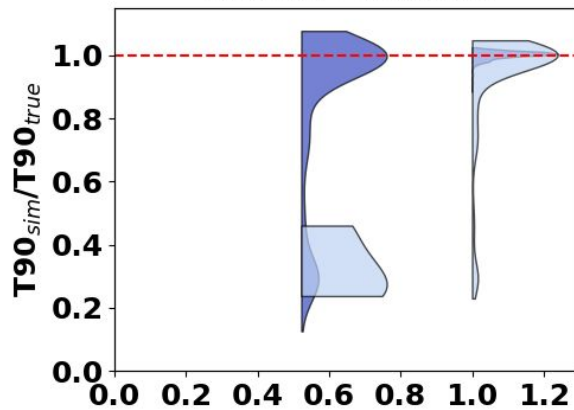
PCODE/INCIDENT ANGLE

# GRB160314A: Short or Long?

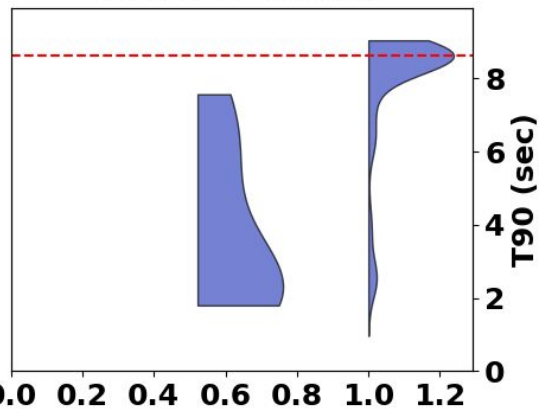


Background (c/s):

1000.0

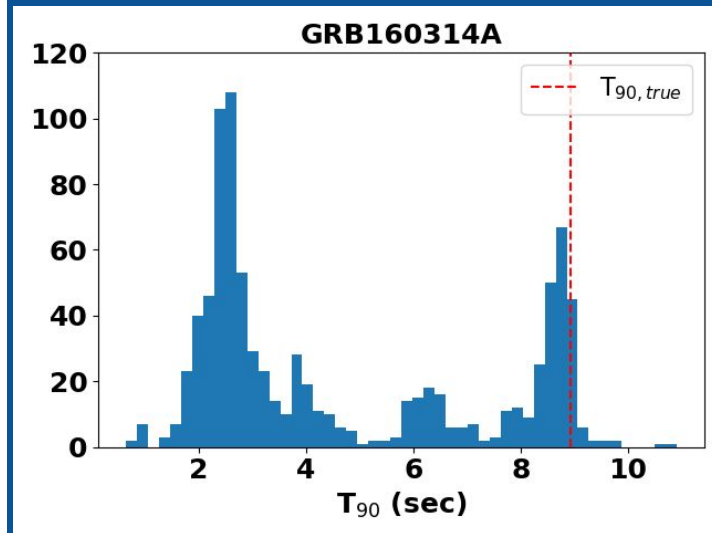
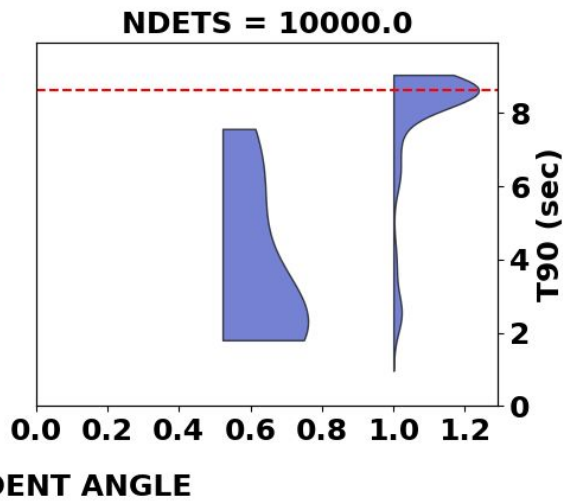
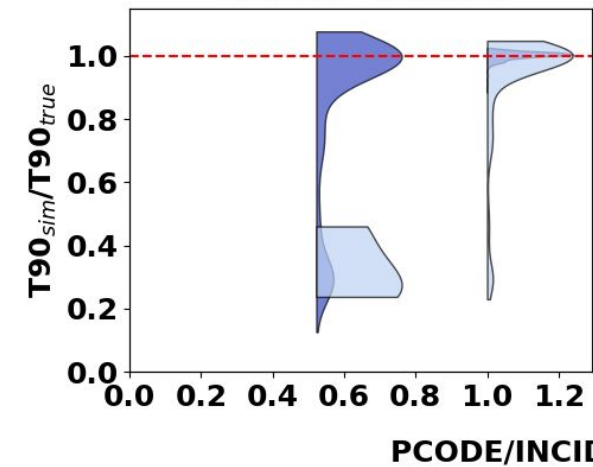
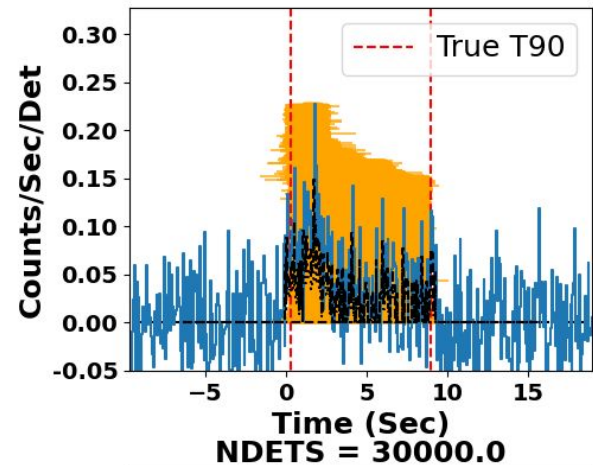


NDETS = 10000.0





# GRB160314A: Short or Long?



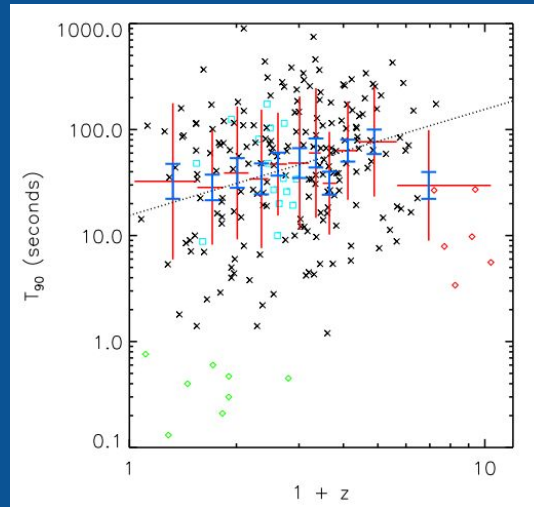
GRB Name	$f(\text{measurable})$	$f(\text{consistent}, 3\sigma)$	$f(\text{consistent}, 1\sigma)$	$T_{90,\text{true}}$ (sec)	Ave. $T_{90,\text{sim}}$ (sec)	90% CI (sec) 68% CI (sec)
(1)	(2)	(3)	(4)	(5)	(6)	(7)
FRED1 ( $1.44 \times 10^{-5}$ erg/cm <sup>2</sup> )	0.914	0.826	0.271	14.14	11.62	[4.008, 15.108] [6.008, 14.108]
FRED2 ( $6.07 \times 10^{-6}$ erg/cm <sup>2</sup> )	0.787	0.828	0.329	14.13	10.86	[5.008, 14.108] [6.008, 14.108]
FRED3 ( $3.64 \times 10^{-6}$ erg/cm <sup>2</sup> )	0.709	0.618	0.259	14.17	10.22	[4.008, 15.108] [5.008, 14.108]
FRED4 ( $1.54 \times 10^{-6}$ erg/cm <sup>2</sup> )	0.571	0.468	0.224	14.14	8.64	[4.008, 14.108] [5.008, 12.108]
GRB160314A	0.289	0.440	0.346	8.64	6.95	[0.942, 9.442] [2.042, 9.042]
GRB150314A	0.990	0.120	0.069	16.0	11.14	[4.003, 19.103] [6.003, 15.103]
GRB120119A	0.911	0.182	0.074	83.0	47.41	[10.014, 107.314] [11.014, 89.214]
GRB110422A	0.999	1.000	0.880	25.0	24.35	[17.033, 27.133] [24.033, 26.133]
GRB071010B	0.799	0.516	0.472	36.0	20.988	[4.007, 39.208] [5.008, 38.208]
GRB051111	0.704	0.472	0.383	65.0	41.491	[7.017, 77.317] [8.017, 75.317]
GRB050219A	0.811	1.000	0.767	24.0	21.787	[13.013, 25.113] [19.013, 24.113]
GRB150314A (no dim-tail)	0.989	1.000	0.260	11.0	11.149	[4.004, 12.104] [11.004, 12.104]
GRB090510	0.652	0.432	0.402	5.69	2.70	[0.064, 6.064] [0.164, 5.864]

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# Distance Considerations

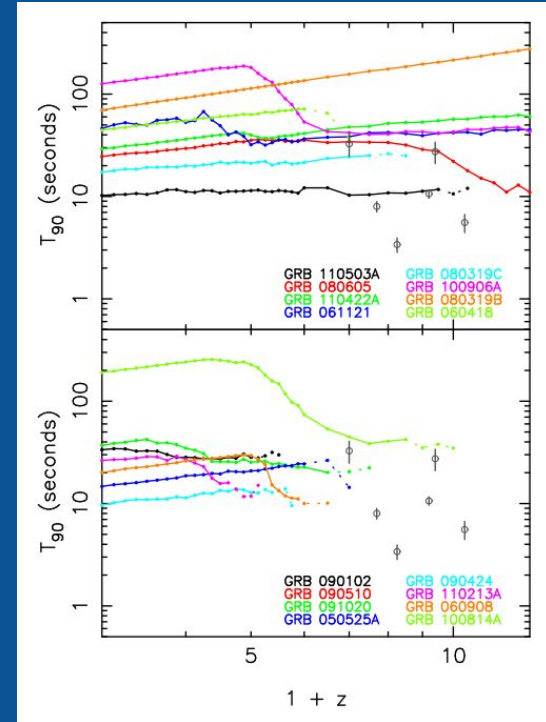
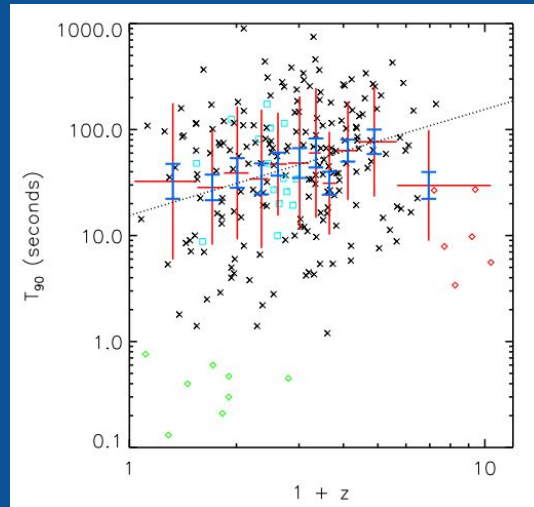


# Distance Considerations



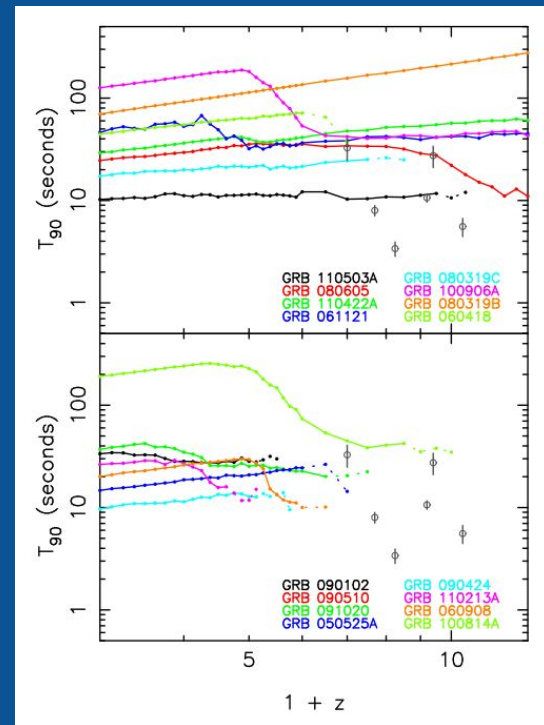
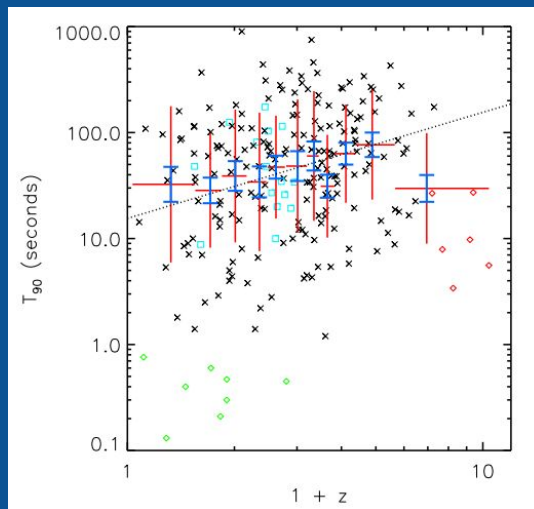
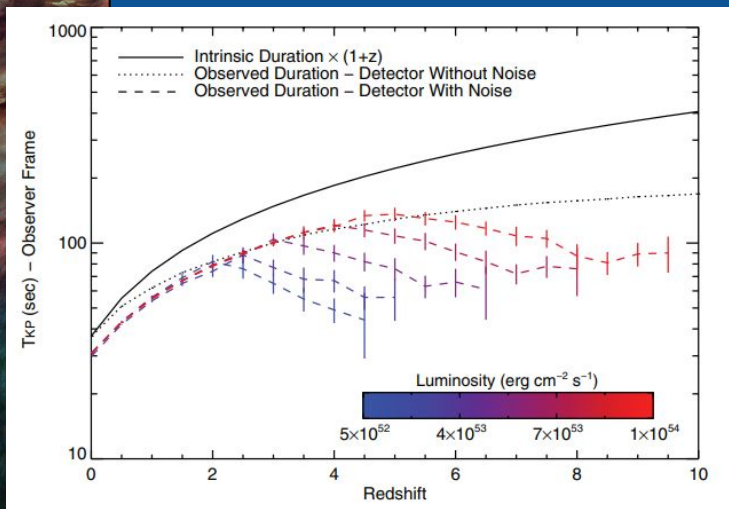
Littlejohns et al., 2013

# Distance Considerations

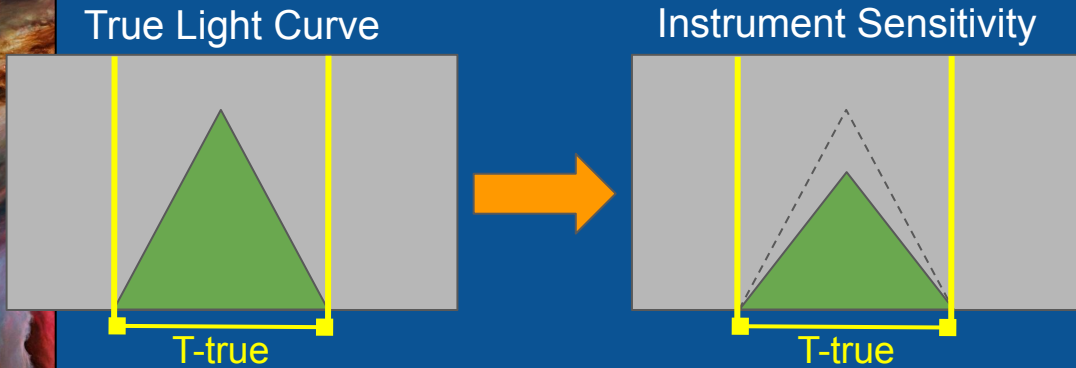


Littlejohns et al., 2013

# Distance Considerations

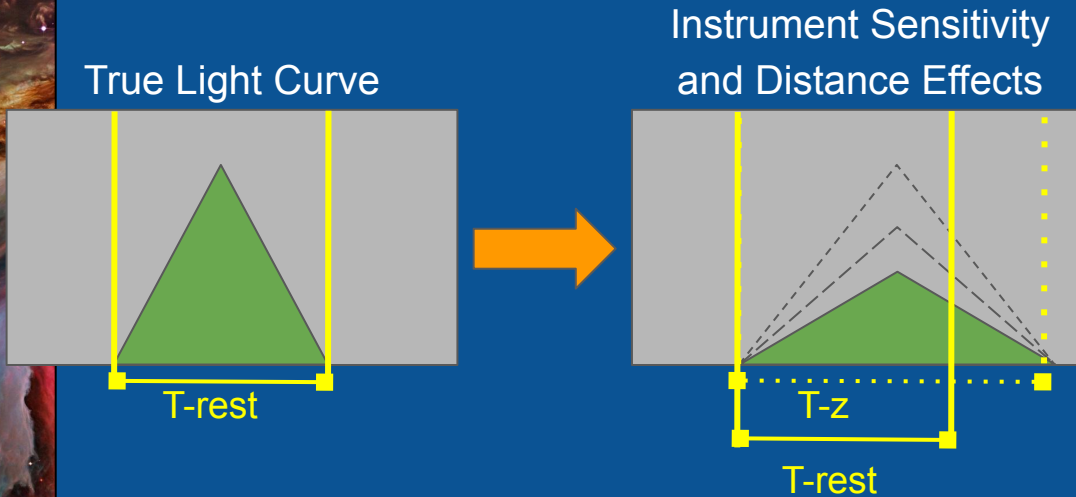


# Distance Effects on GRB Light Curves

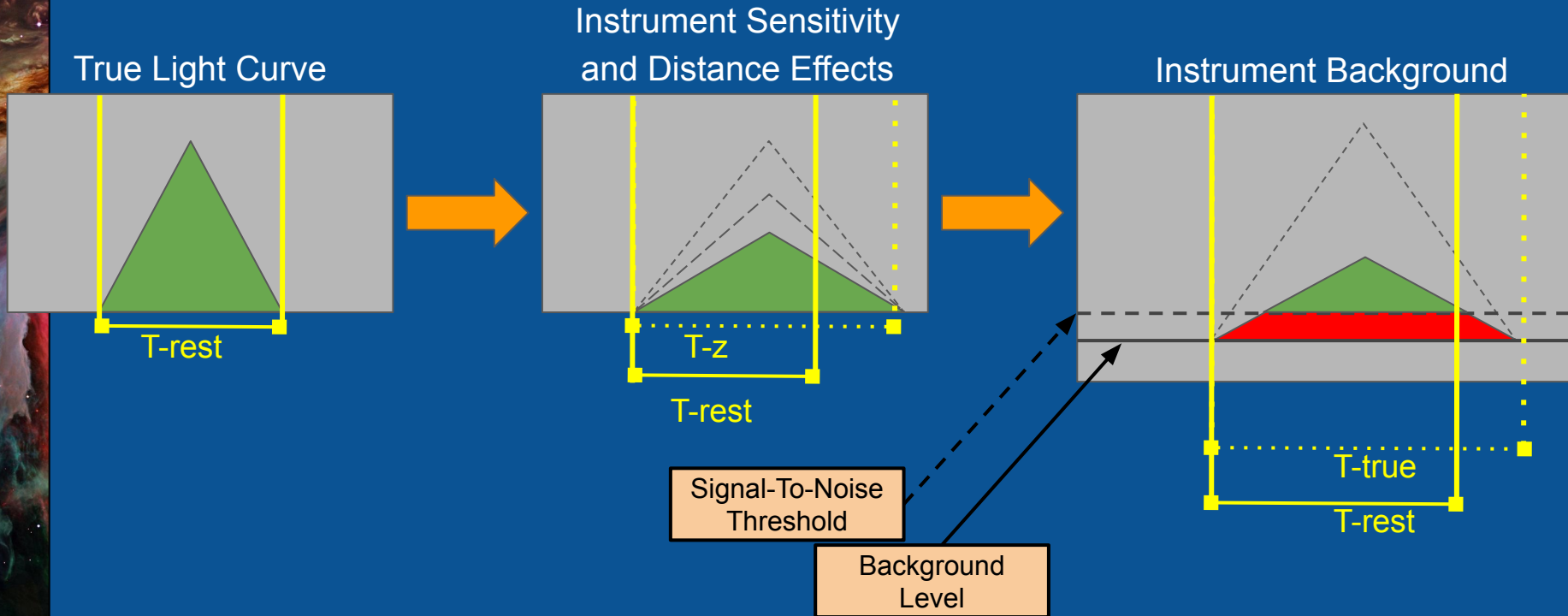




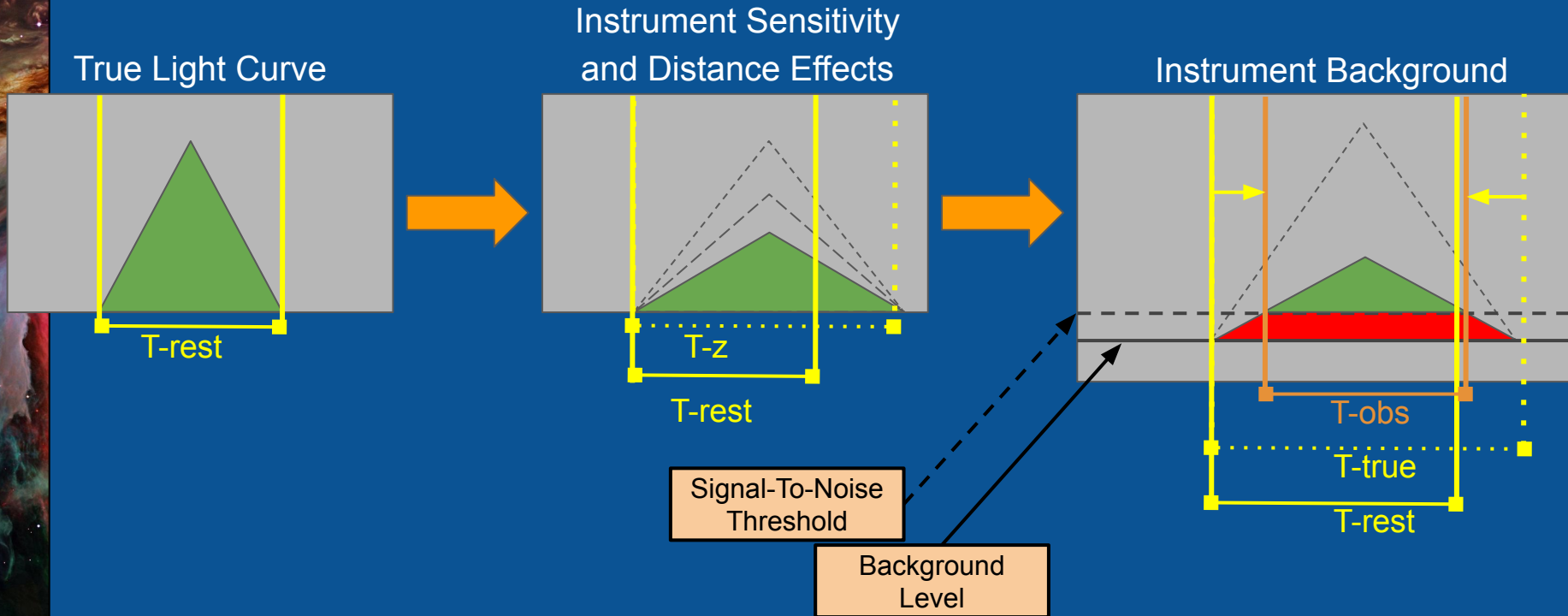
# Distance Effects on GRB Light Curves



# Distance Effects on GRB Light Curves

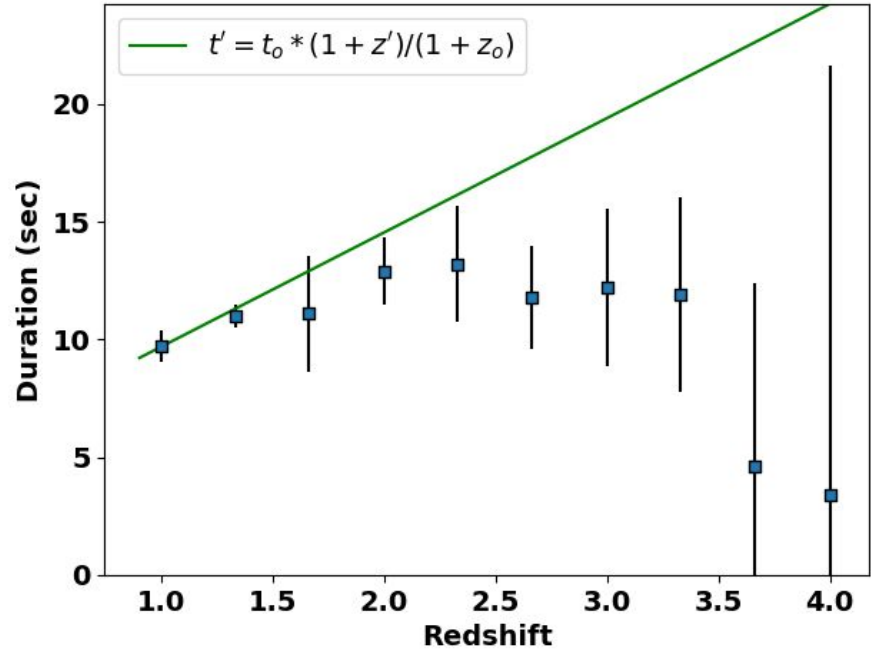
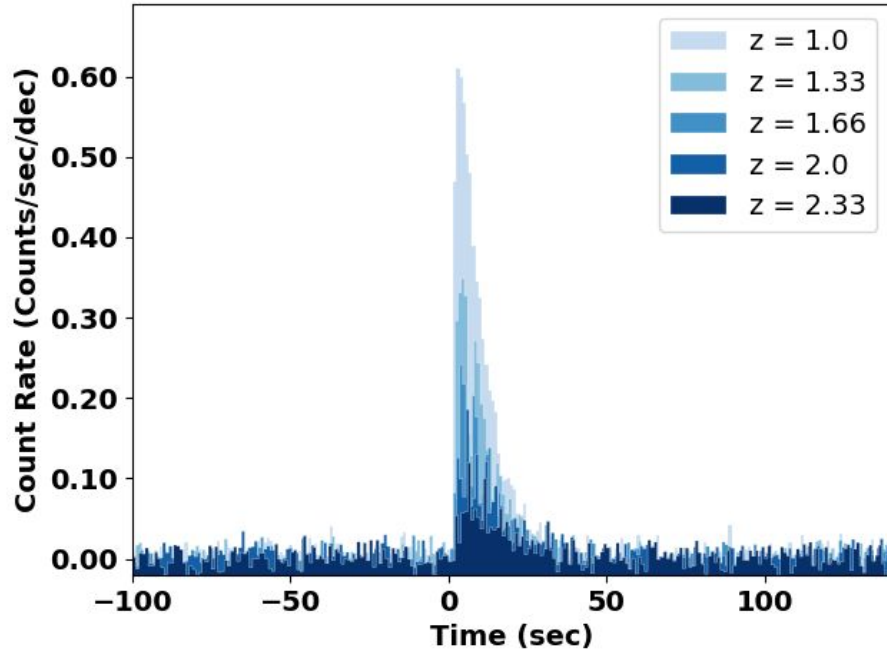


# Distance Effects on GRB Light Curves



# Distance Considerations - Results

This is preliminary, to be further investigated at a later time.

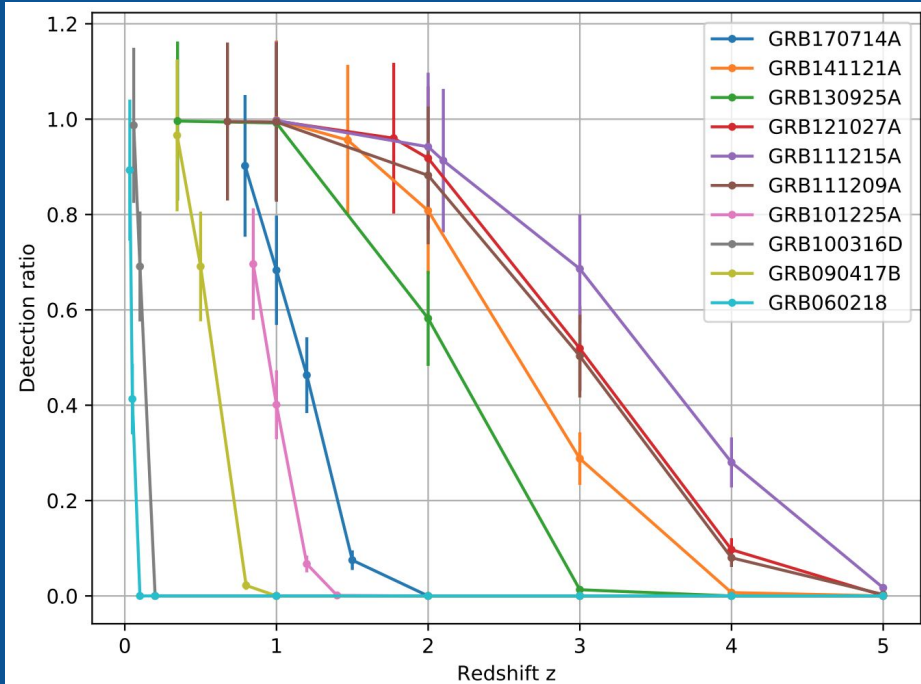
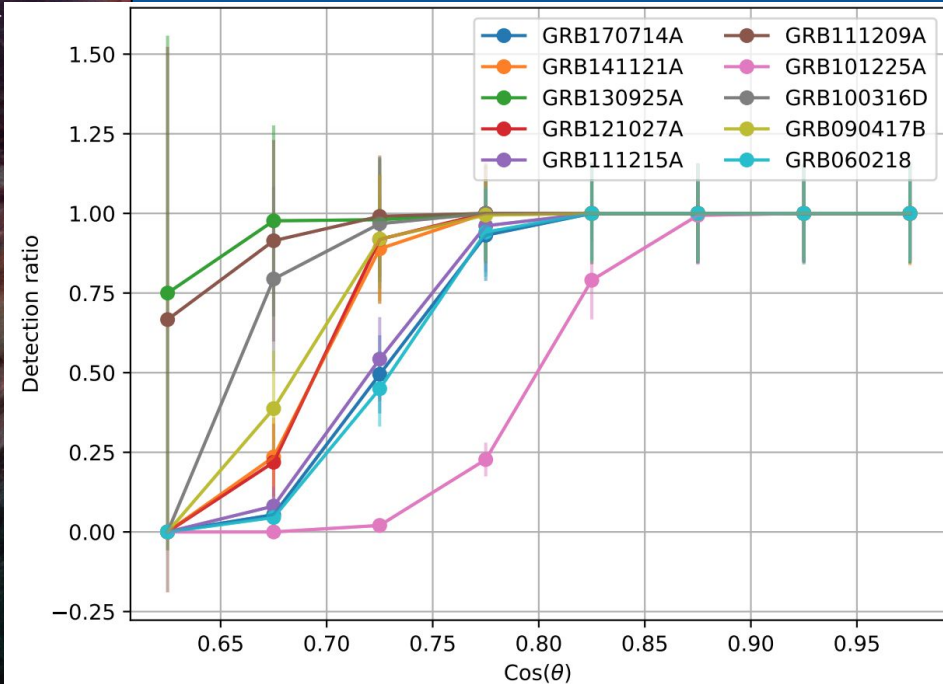


A vertical strip on the left side of the slide shows a colorful cosmic nebula with swirling clouds of gas in shades of red, orange, and green, set against a dark background with some stars.

# Conclusions

- 1. Swift/BAT GRB duration measurements are highly impacted by observing conditions**
- 2. As instrument sensitivity decreases**
  - **Average T90 becomes shorter and**
  - **T90 uncertainty becomes larger**
3. The PCODE has the strongest influence
4. Strong dependence on light curve
5. Most of our sample, consistent with intrinsic durations in only ~20% - 40% of simulations.
6. Light curves with a intrinsic T90 > 2 sec may be observed with T90 < 2 sec

# Instrumental Considerations - SVOM/ECLAIRs



Dagoneau, Schanne, Atteia, Götz, and Cordier (2020)

See also B. Arcier, et al. 2021 for SVOM/ECLAIRs detection capability of short high-energy transients

# Next Steps

- Include complex light curves
- Cosmological distance effects
- Investigate bias in the observed luminosity and T90 distributions
- Apply to other instruments



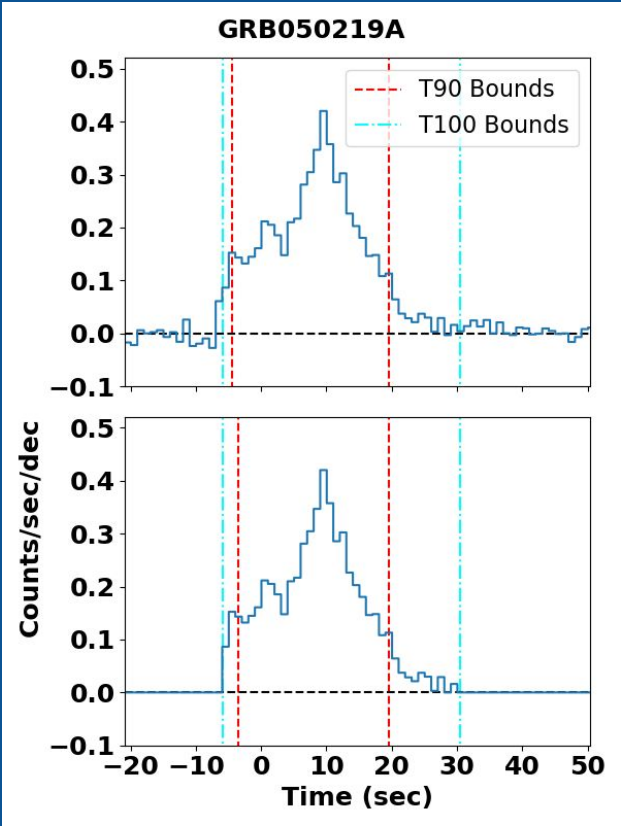
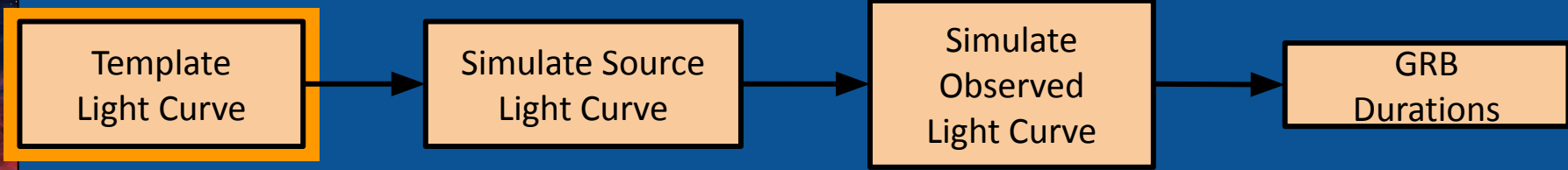


Thank you!



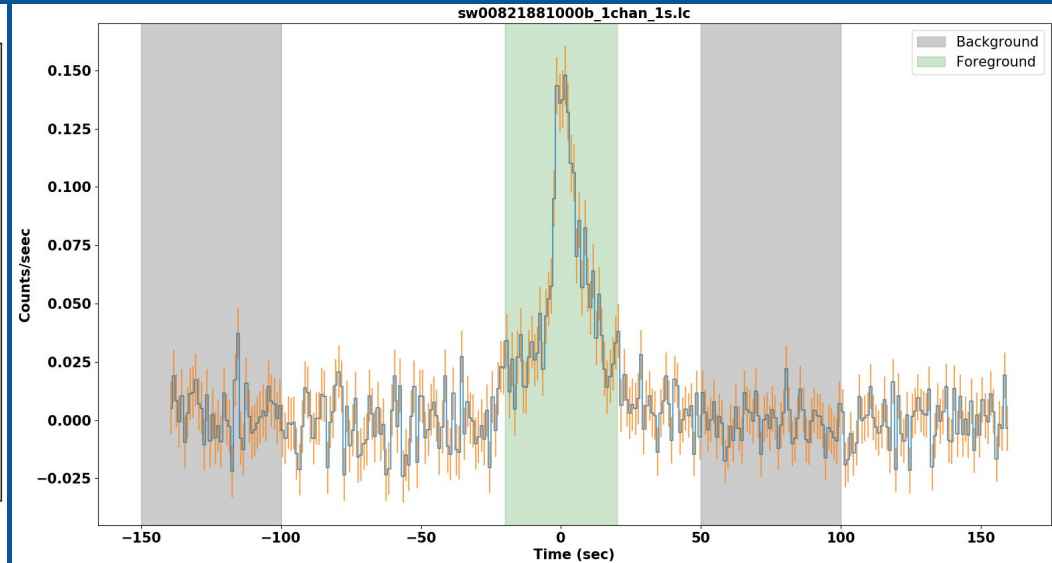
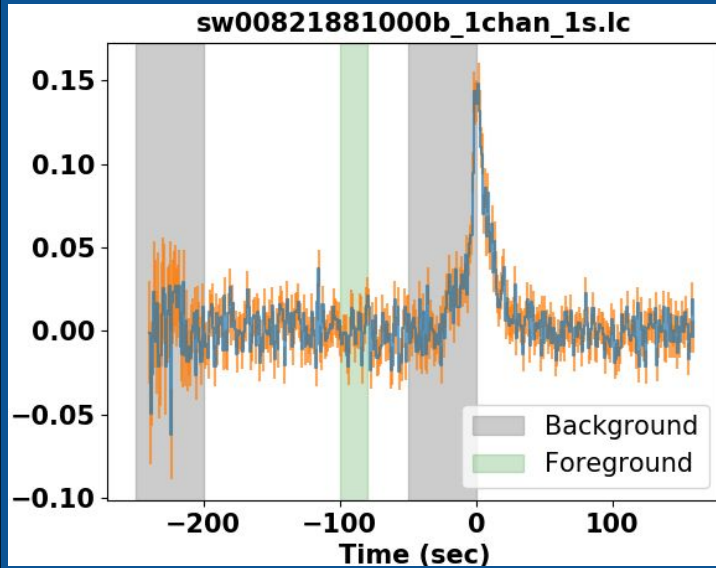


# Backup Slides

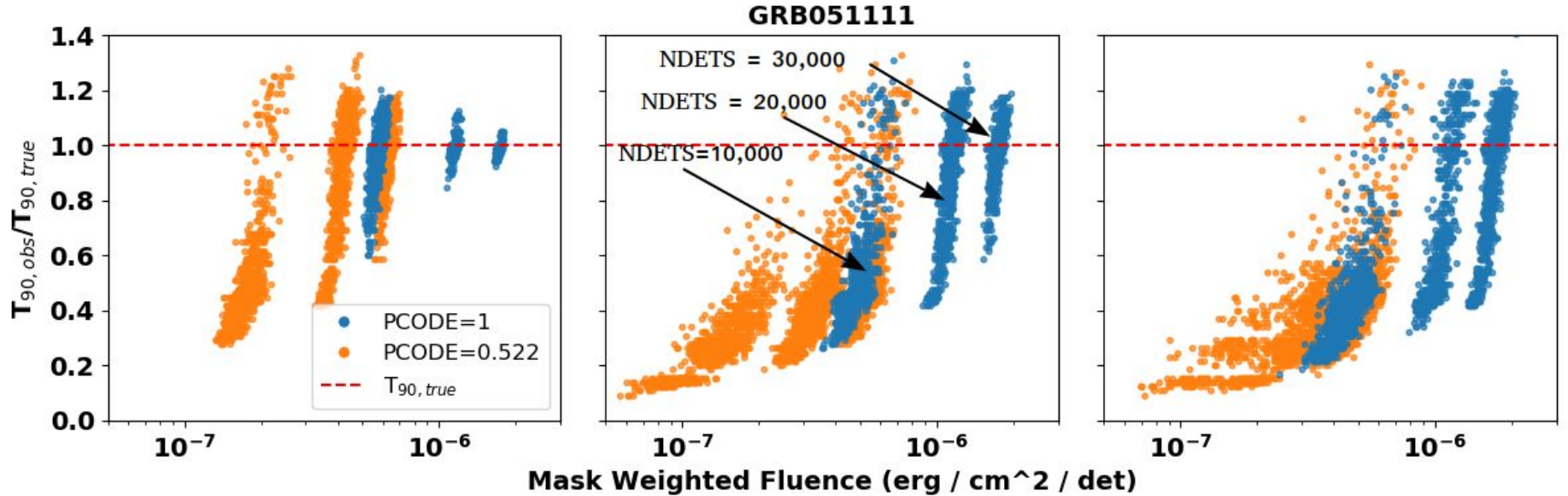


# Bayesian Blocks

Few hundred sliding Bayesian Blocks combinations



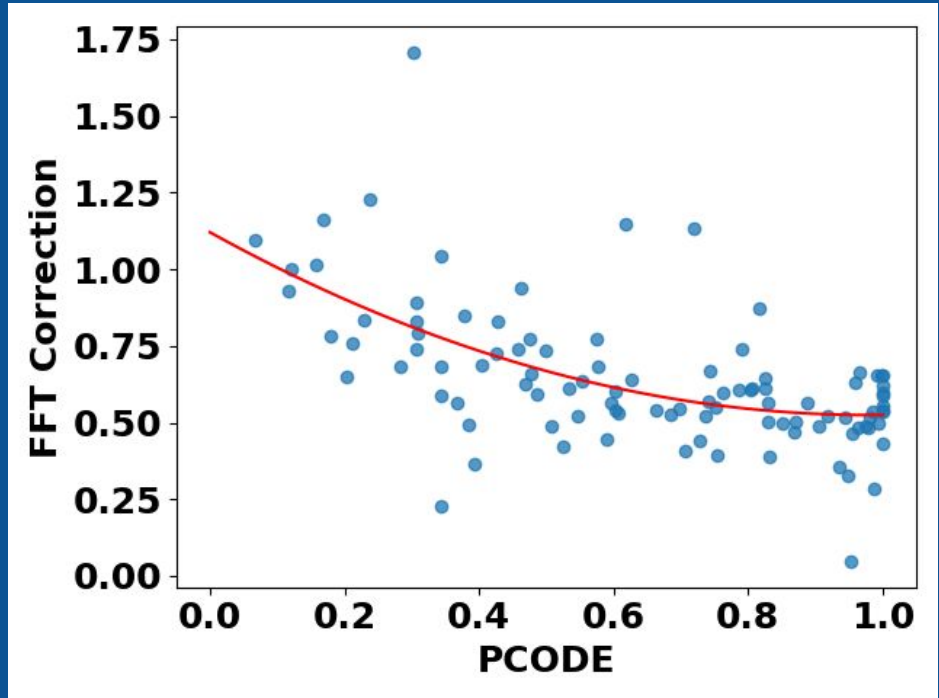
# Signal-to-Noise Ratio



# FFT Loss Factor

FFT causes additional loss of signal

We calculated the loss for 100 GRBs and Fit a line as a function of PCODE.

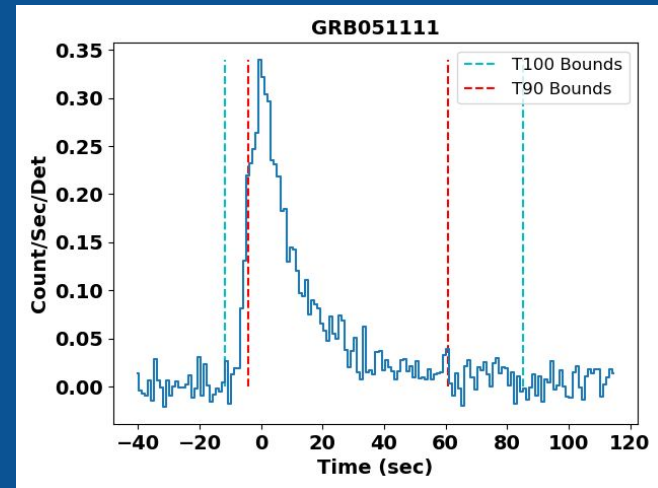
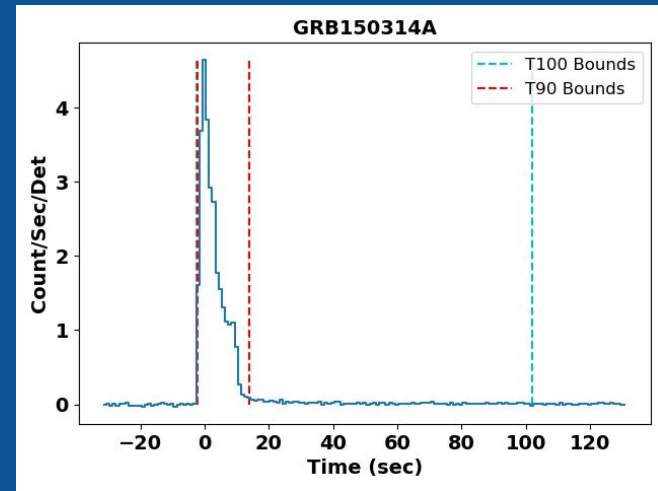
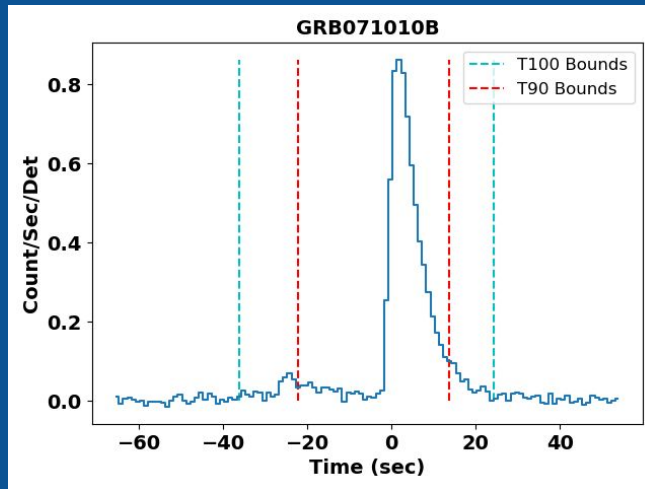


# Light Curve Sample

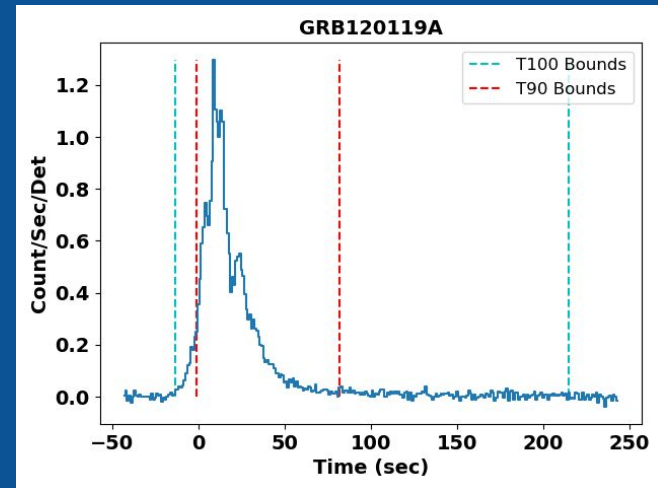
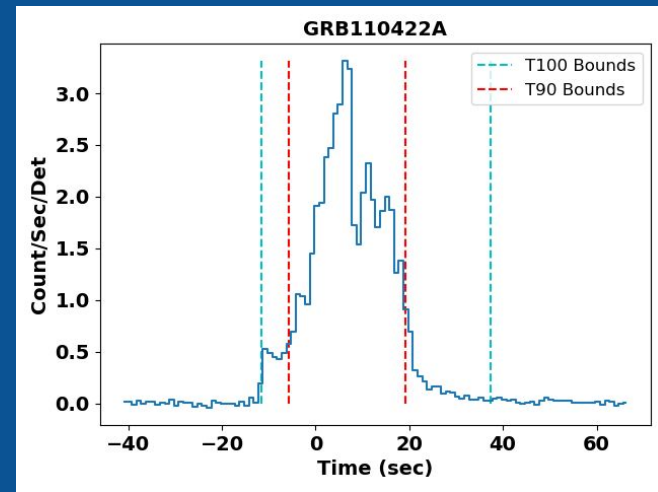
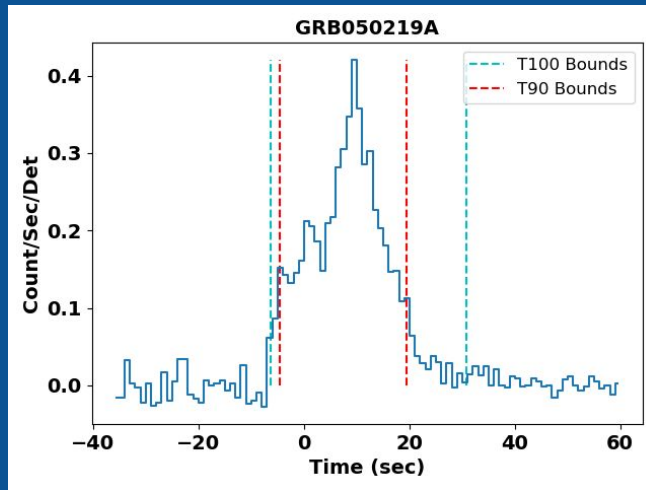
GRB Name	$z$	$T_{90}$ (sec)	Fluence (erg/cm <sup>2</sup> )	$\alpha$	PCODE	$\theta_{\text{inc}}$	SNR	Description
(1)	(2)	(3)	(4)	(5)	(6)	(7)	(8)	(9)
GRB160314A	1.726	8.73	$3.75 \times 10^{-07}$	-1.53	0.75	19.69°	14.18	Short pulse
GRB150314A	1.76	14.8	$5.13 \times 10^{-05}$	-1.08	0.344	35.1°	256	FRED-like with dim tail
GRB120119A	1.73	68.0	$3.17 \times 10^{-05}$	-1.38	1.02	5.13°	45.46	Symmetric-like
GRB110422A	1.77	25.8	$5.56 \times 10^{-05}$	-0.831	0.227	44.7°	27.95	Symmetric-like
GRB090510	0.903	5.664	$1.46 \times 10^{-06}$	-1.06	0.162	46.07°	145.49	Short-hard spike with soft tail
GRB071010B	0.947	36.124	$6.21 \times 10^{-06}$	-1.97	0.8438	29.04°	52.96	FRED-like with dim pre-trigger emission
GRB051111	1.55	64.0	$7.94 \times 10^{-06}$	-1.32	0.594	27.2°	37.09	Broad FRED-like
GRB050219A	0.211	23.8	$4.53 \times 10^{-06}$	-0.124	0.232	43.1°	14.69	Symmetric-like

# GRB Light Curves

## FRED-like



# GRB Light Curves Symmetric





# GRB Light Curves

